

Space – The Final Frontier

Jason Drury

PhD Student

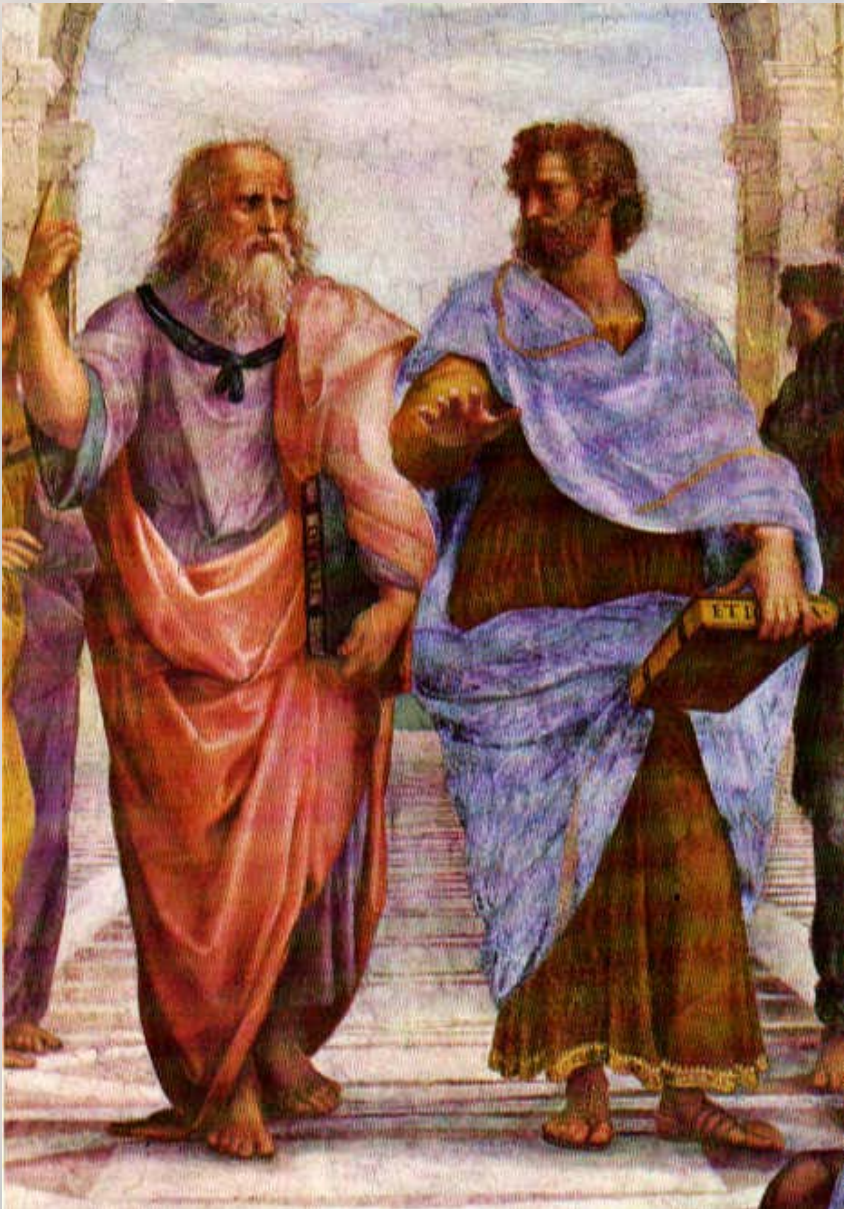
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THE UNIVERSITY OF
SYDNEY

The Fascinating Cosmos

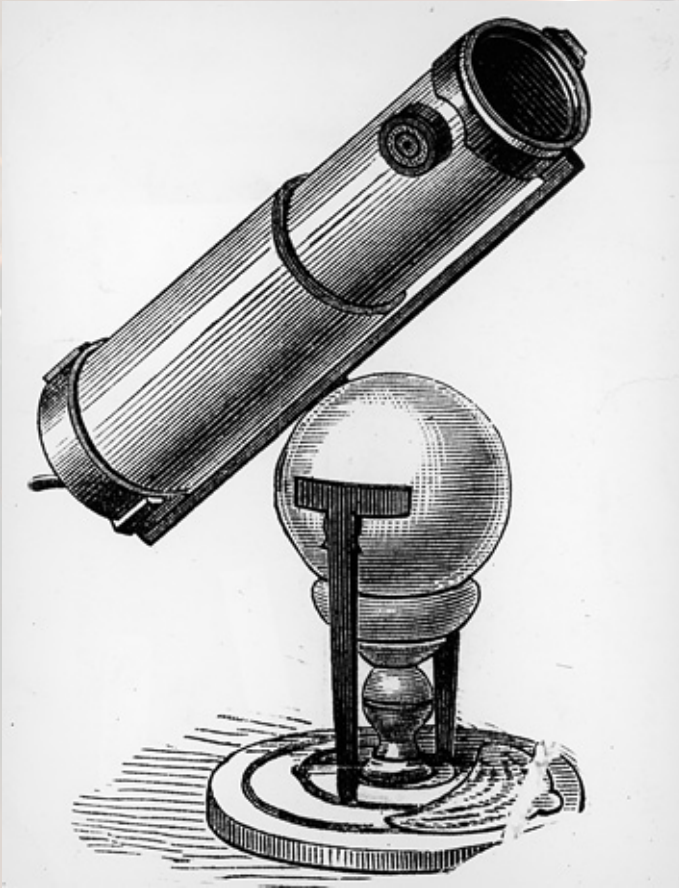


Credit: Zebu.uoregon.edu

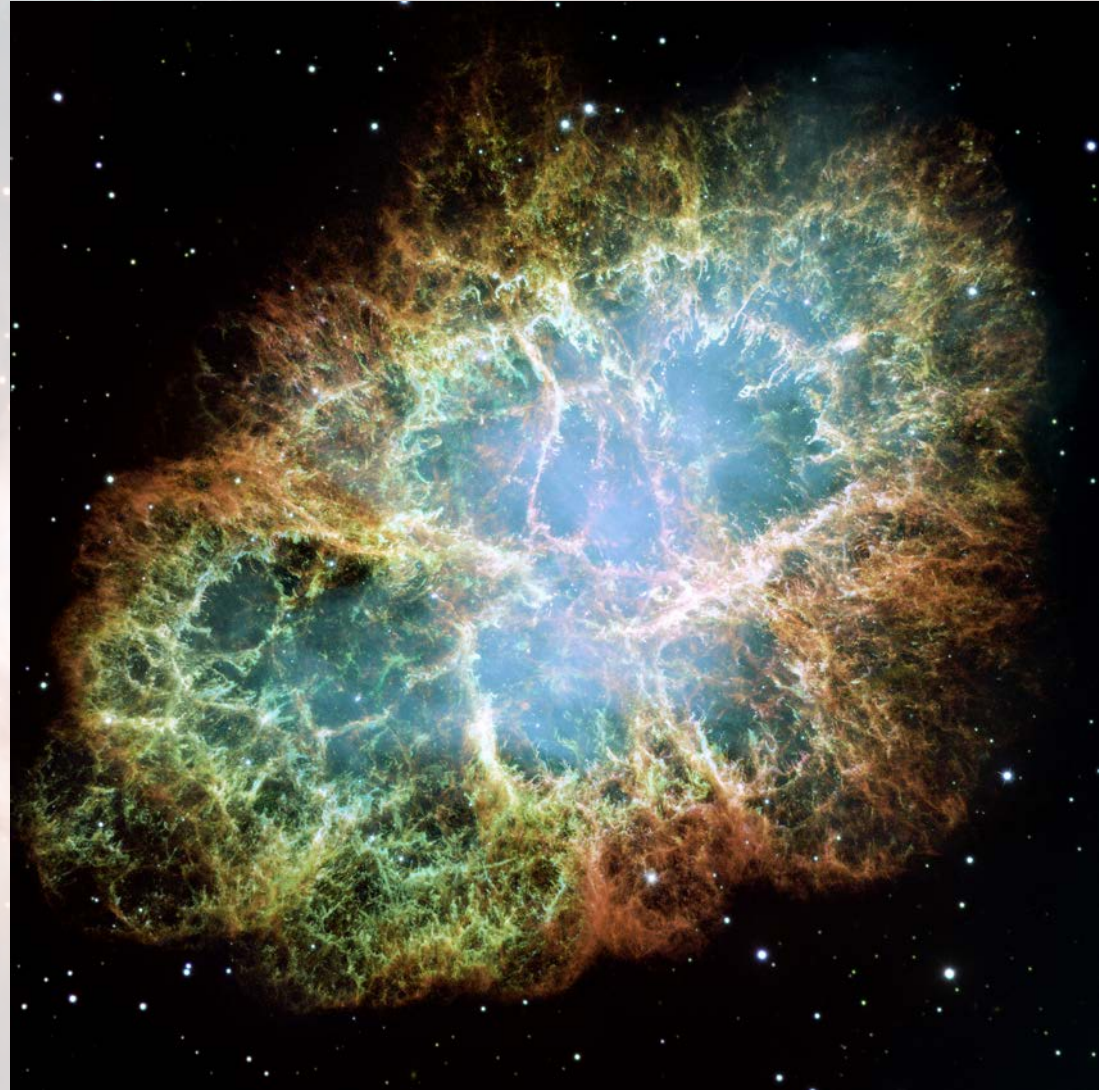


Credit: Grant – Celestial Orbs in the Latin Middle Ages

The Fascinating Cosmos

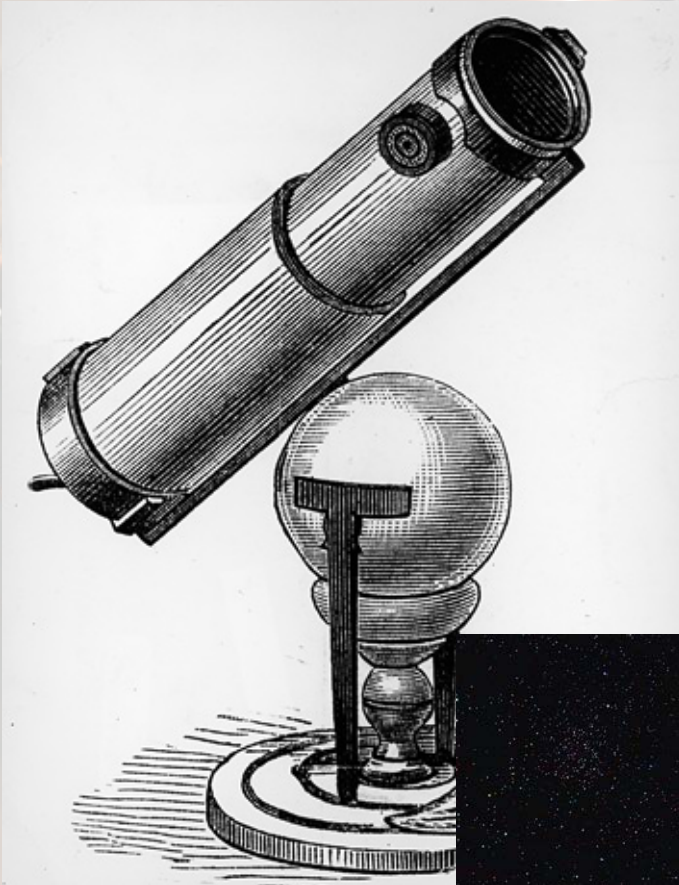


Credit: Hulton
Archive/Getty
Images

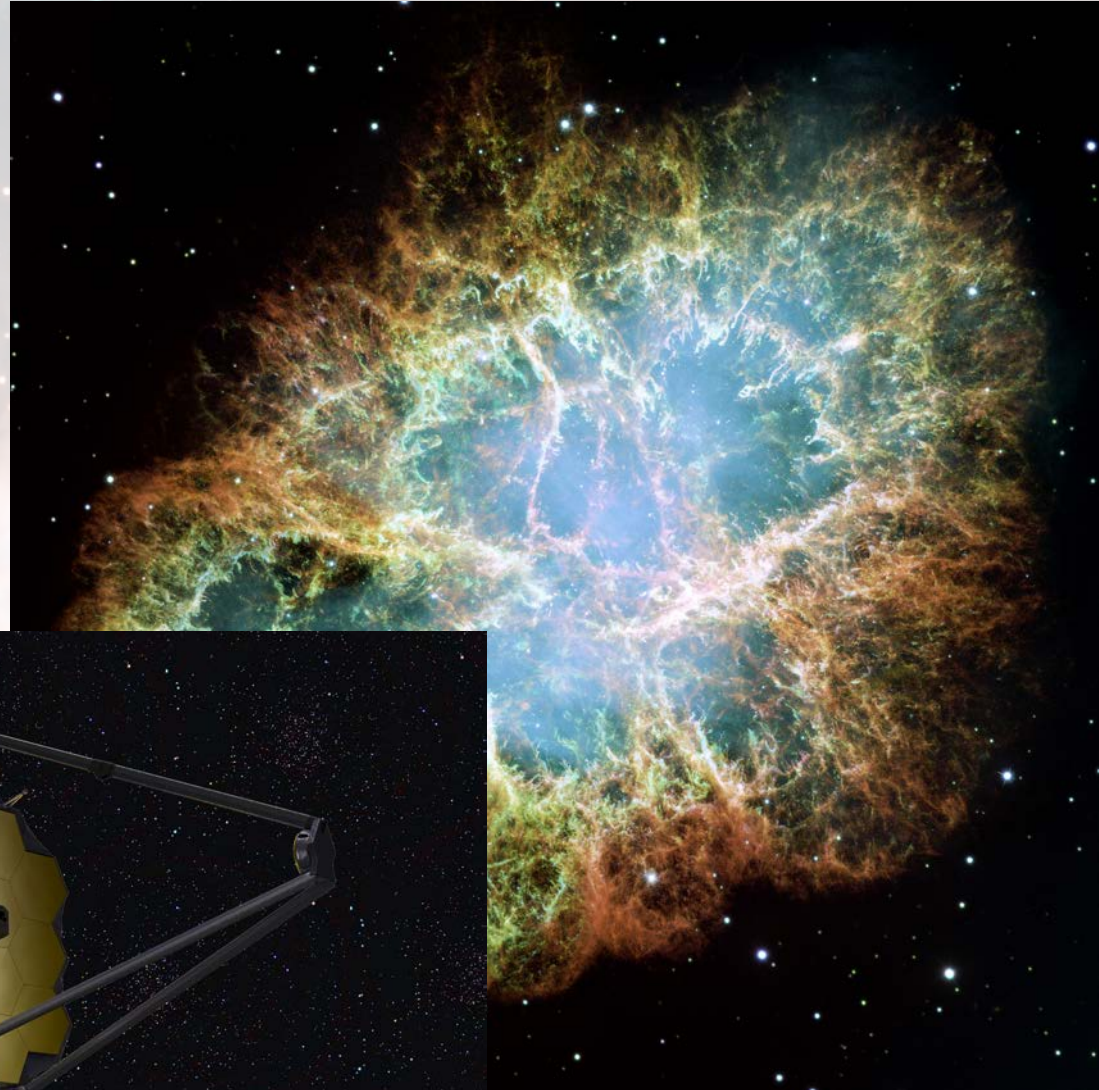


Credit: NASA/Hubble
Telescope

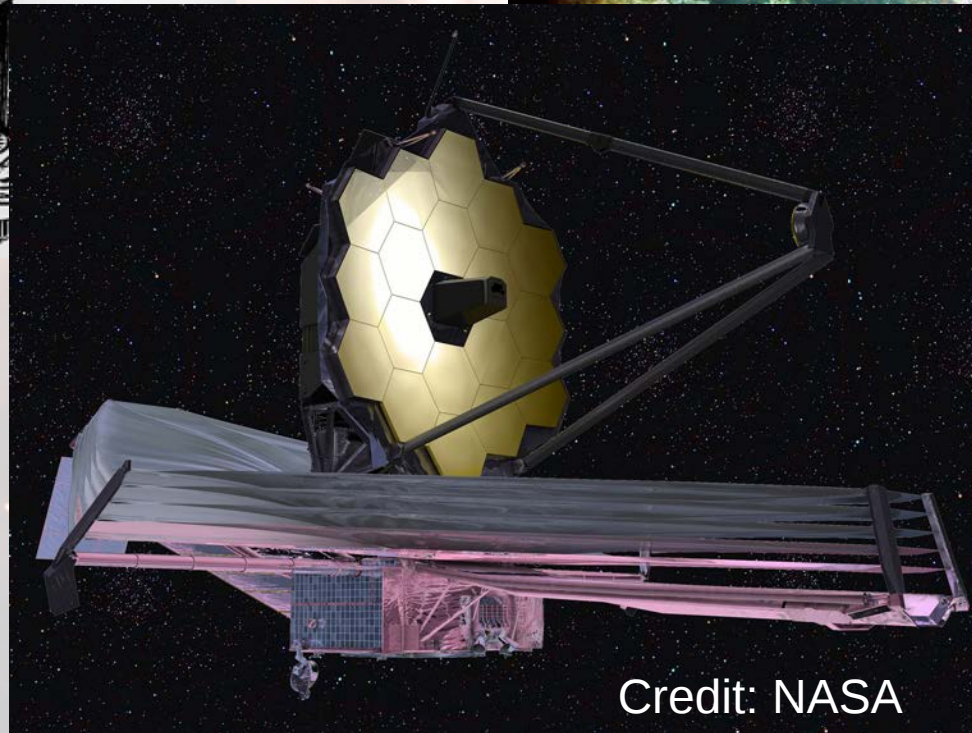
The Fascinating Cosmos



Credit: Hulton
Archive/Getty
Images



Credit: NASA/Hubble
Telescope



Credit: NASA

Newtonian Gravity

Law of Universal Gravitation:

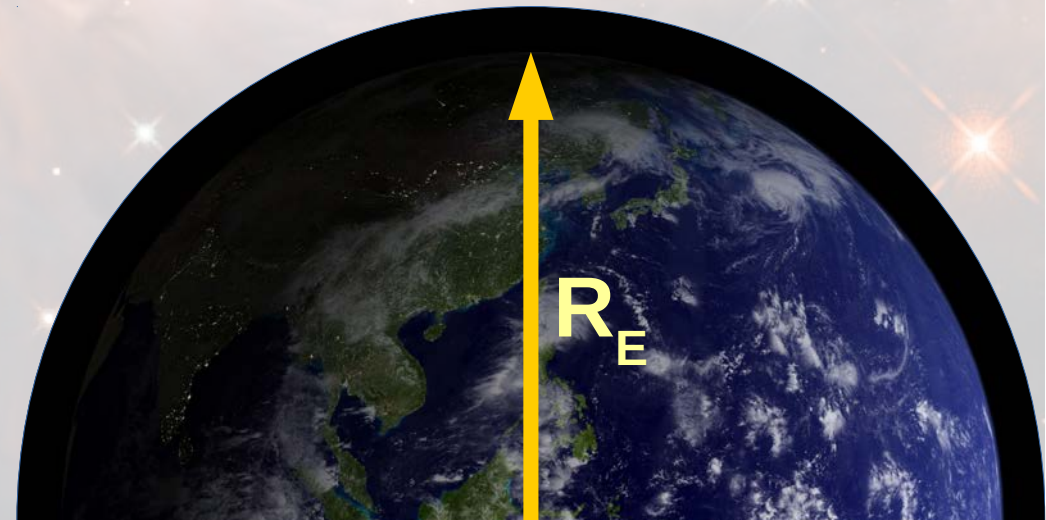
$$F_g = \frac{Gm_1m_2}{d^2}$$

Newtonian Gravity

Law of Universal Gravitation:

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$$F_g = \frac{GM_E}{R_E^2} m$$

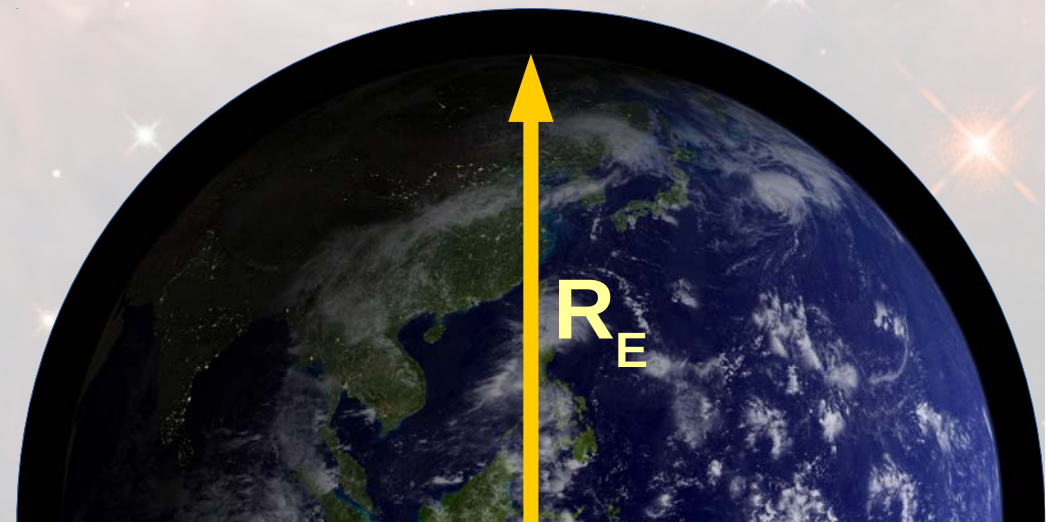


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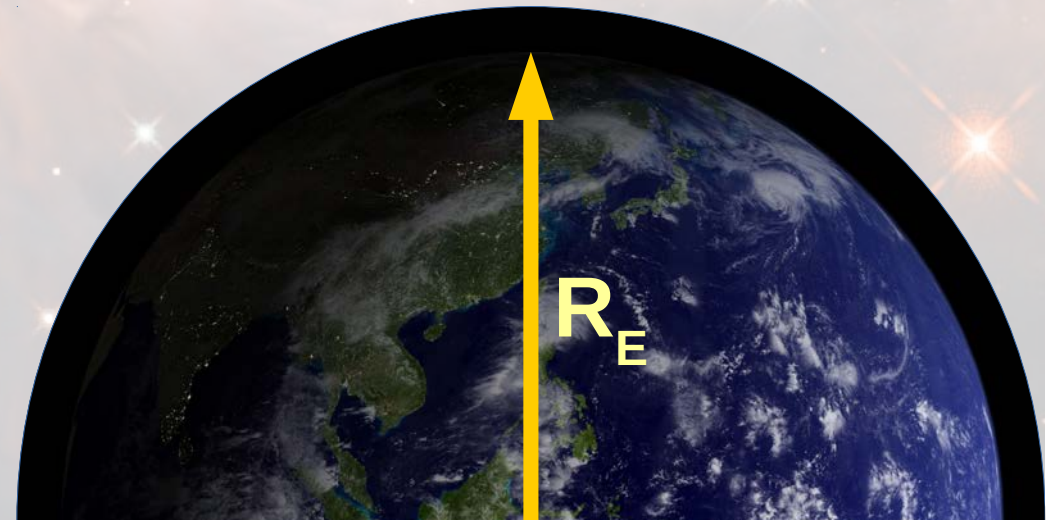


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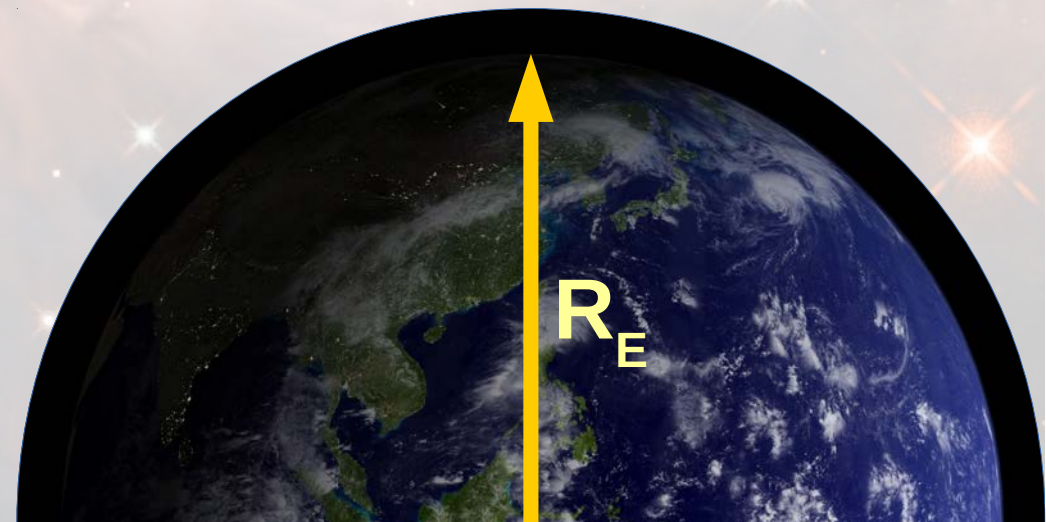
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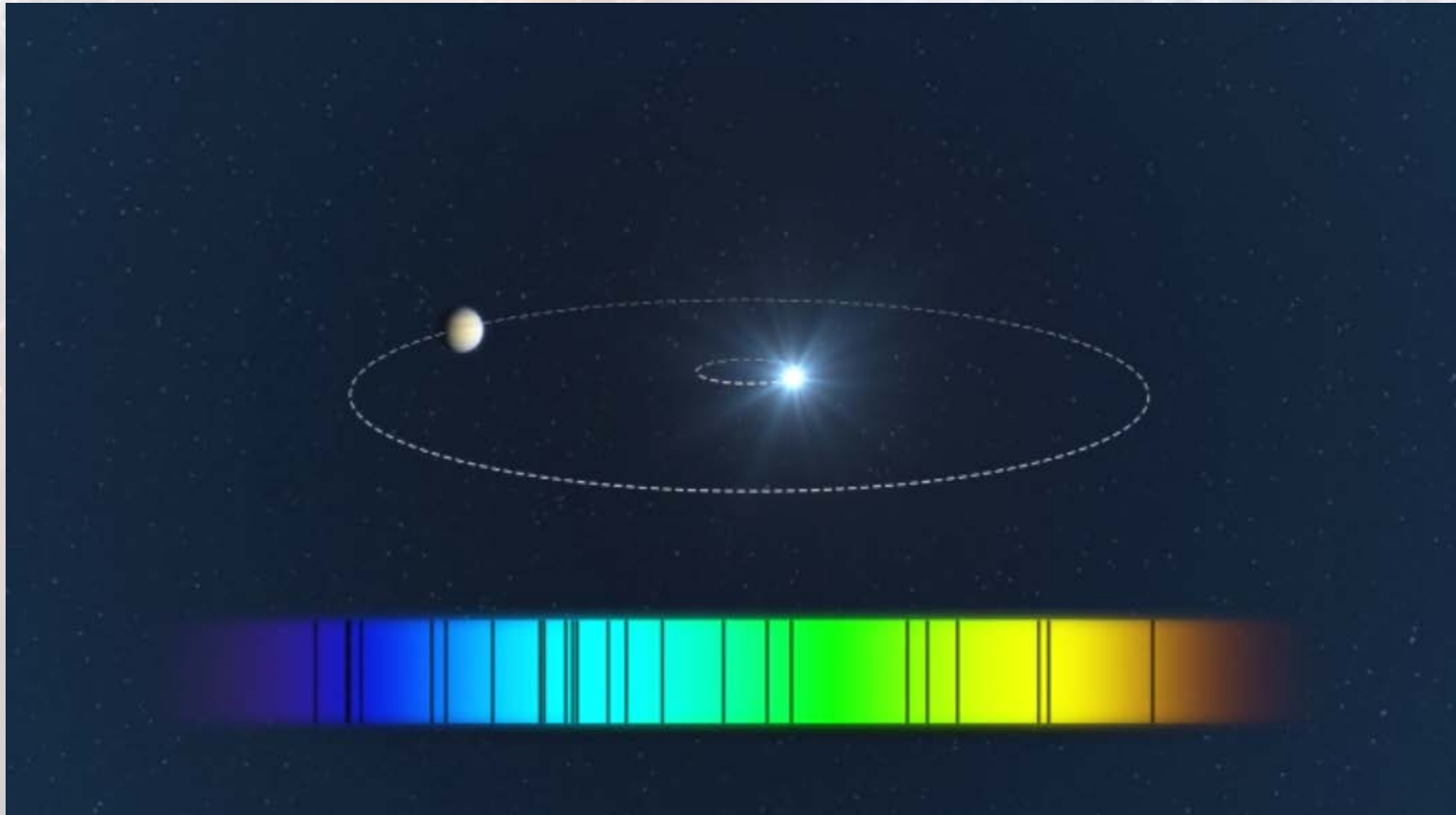
$$g = a = \frac{GM_E}{R_E^2}$$



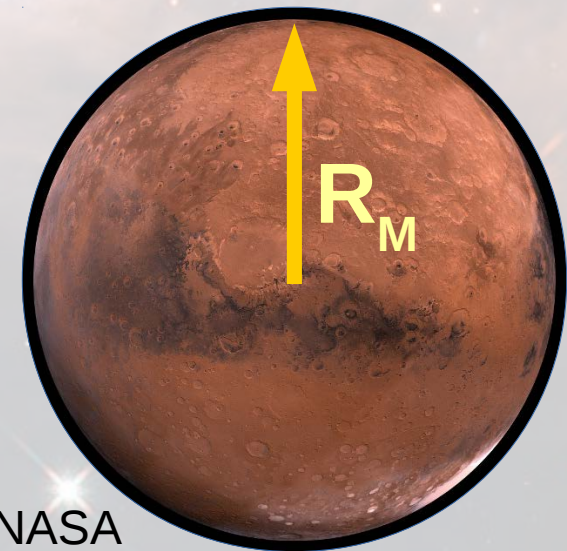
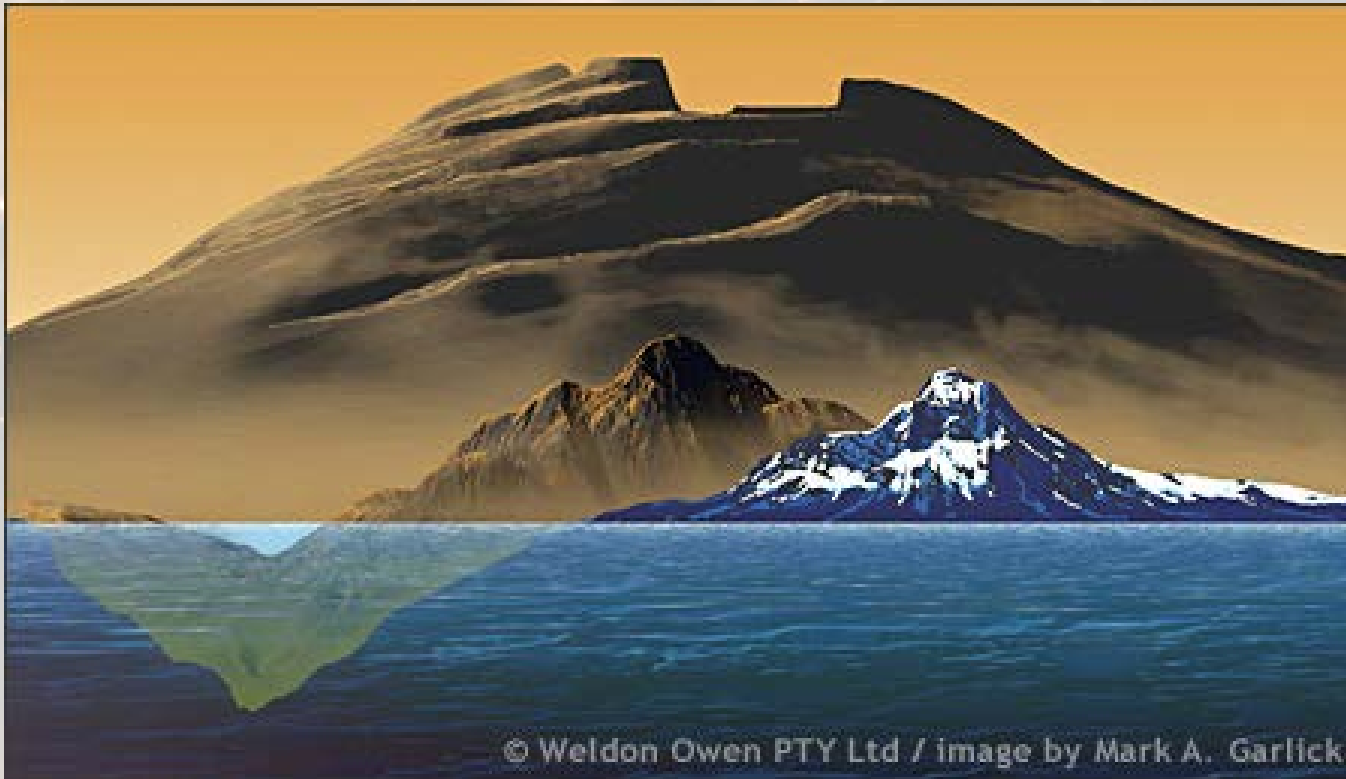
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Newtonian Gravity



Credit: NASA

Newtonian Gravity

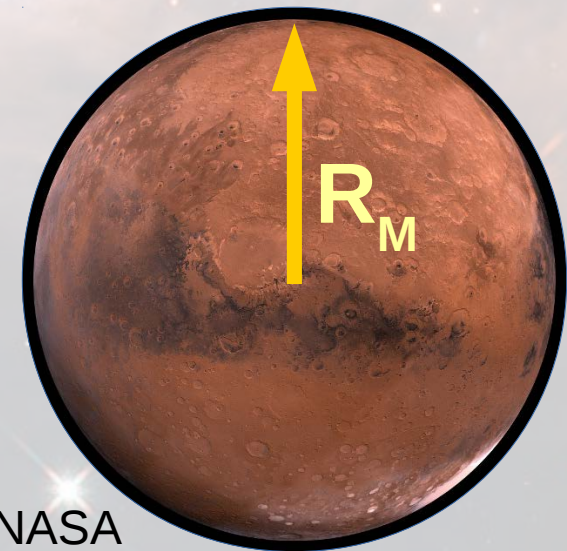
$$g = \frac{GM_{Mars}}{R_{Mars}^2}$$

$$M_{Mars} = 6.42 \times 10^{23} \text{ kg}$$

$$R_{Mars} = 3389.5 \text{ km}$$

$$H_{OM} = 21.2 \text{ km}$$

$$G = 6.67 \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2}$$



Newtonian Gravity

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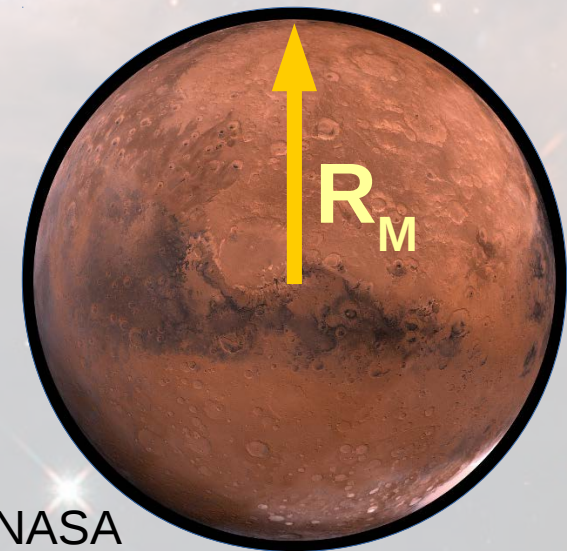
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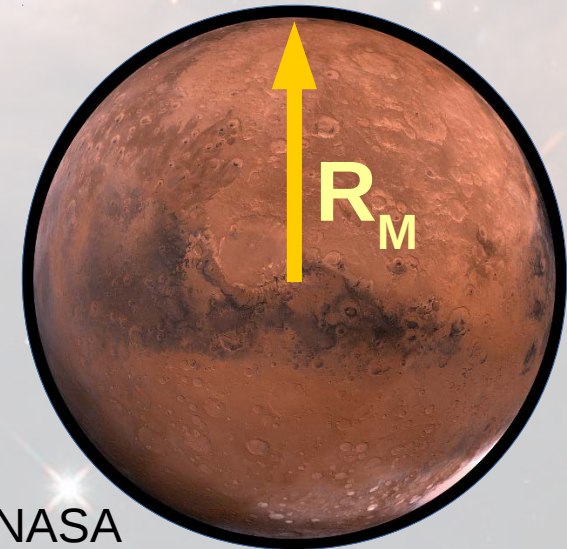
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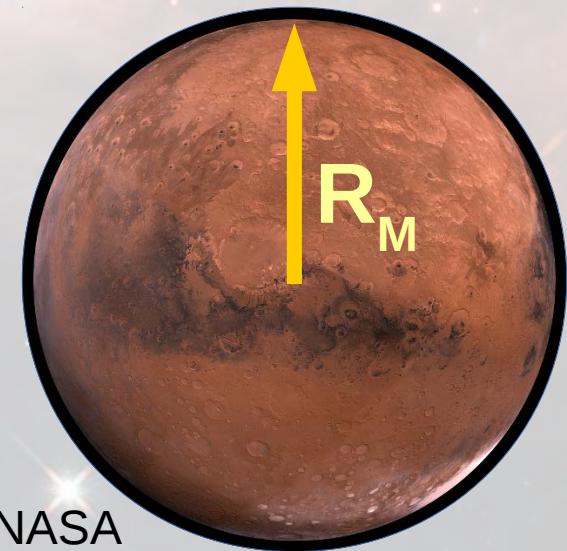
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Sanity Check!!



Newtonian Gravity

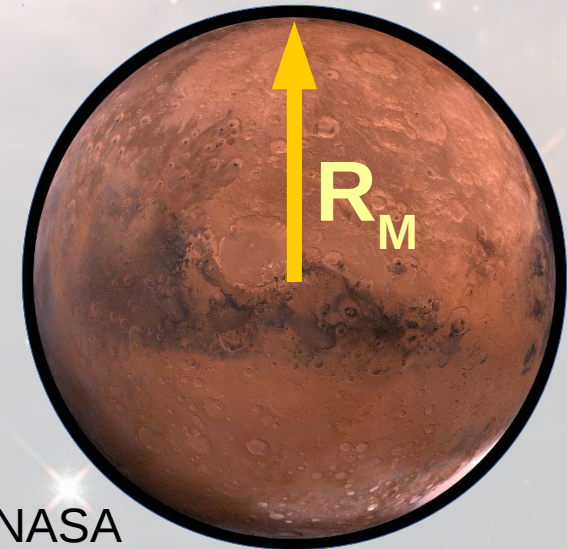
$$g = \frac{GM_{Mars}}{(R_{Mars} + H_{OM})^2}$$

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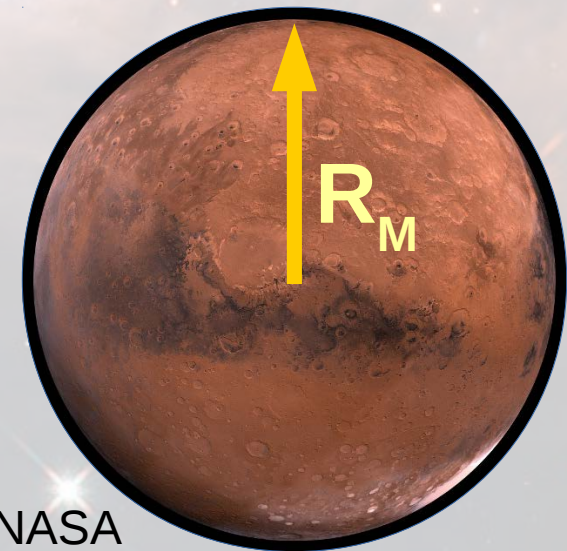
$$g = \frac{6.67 * 10^{-11} * 6.42 * 10^{23}}{(3410.7 * 10^3)^2}$$

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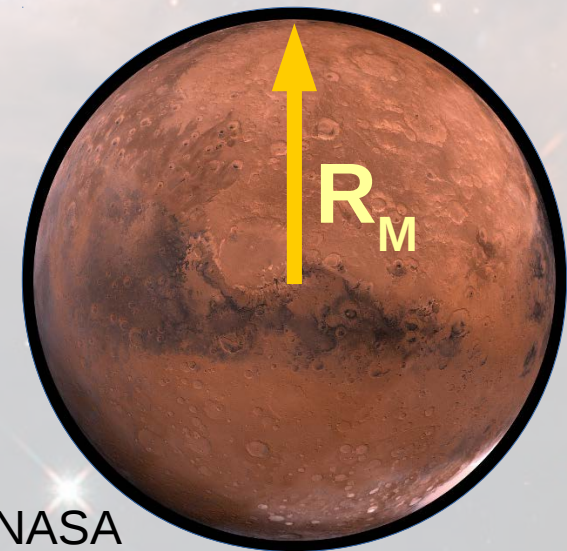
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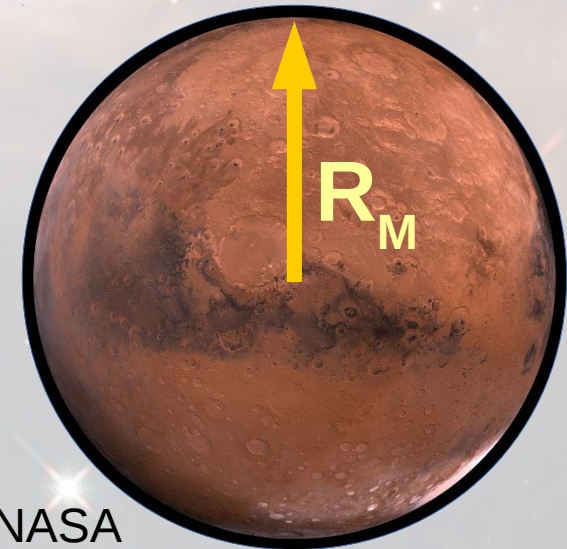
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Significant Figures

$$M_{Mars} = 6.42 \times 10^{23} \text{ kg}$$

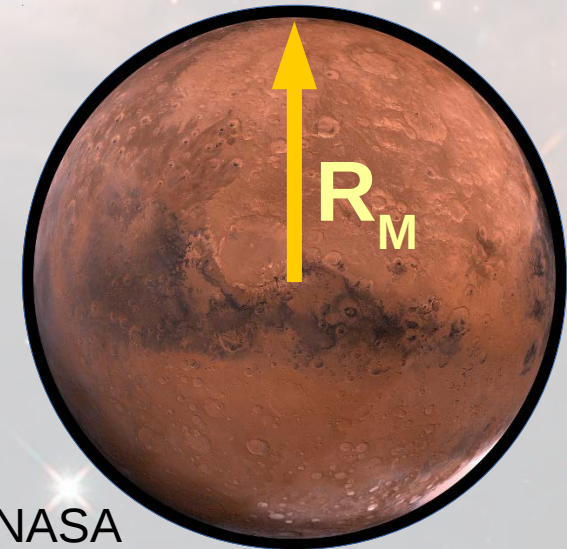
$$R_{Mars} = 3389.5 \text{ km} = 3.3895 \times 10^3 \text{ km} \text{ (5 sig. figs)}$$

$$H_{OM} = 21.2 \text{ km} = 2.12 \times 10^1 \text{ km} \text{ (3 sig. figs)}$$

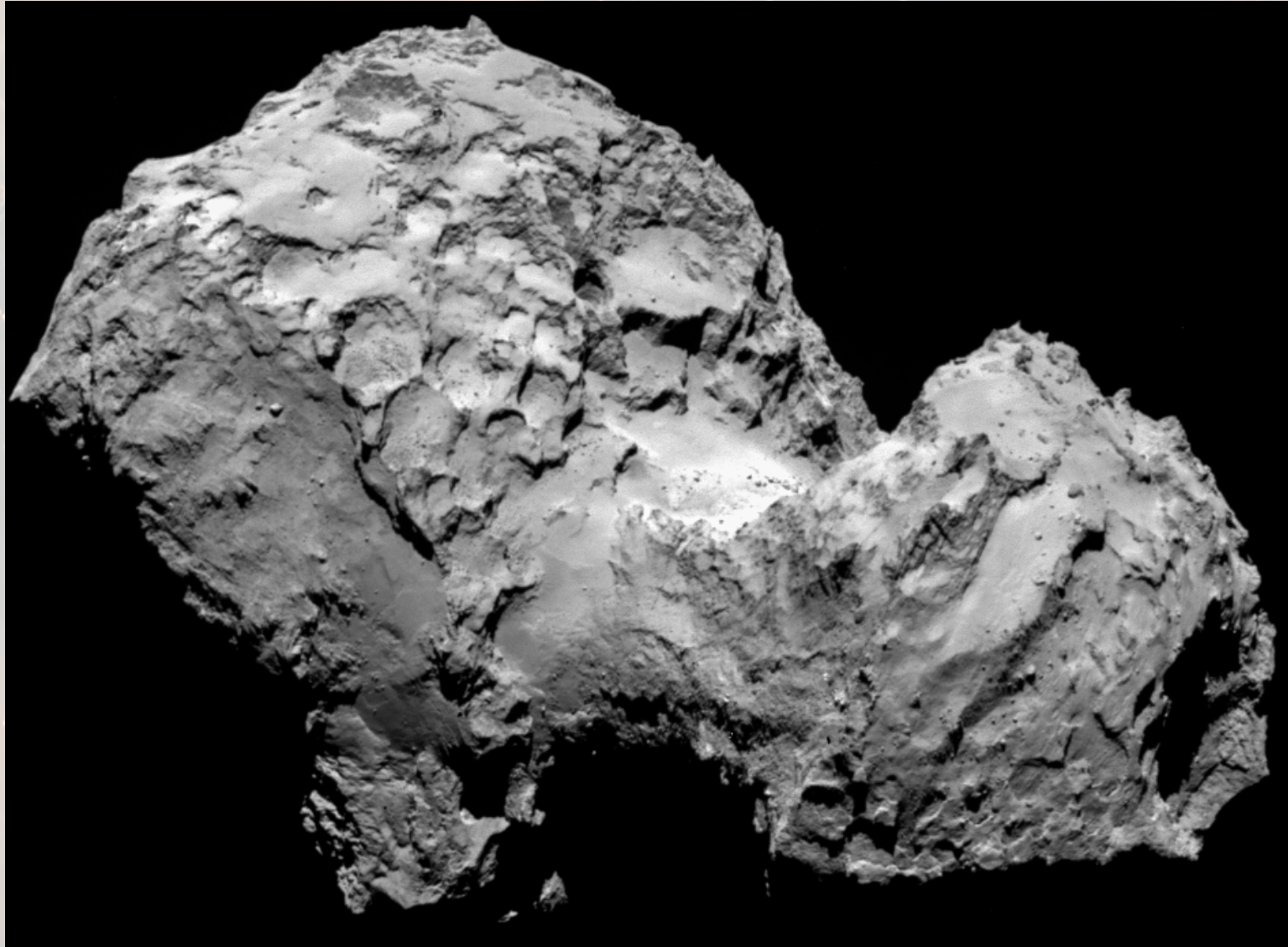
$$G = 6.67 \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2}$$

Another example:

$$0.061 \text{ m} = \underline{6.1} \times 10^{-2} \text{ m} \text{ (2 sig. figs)}$$

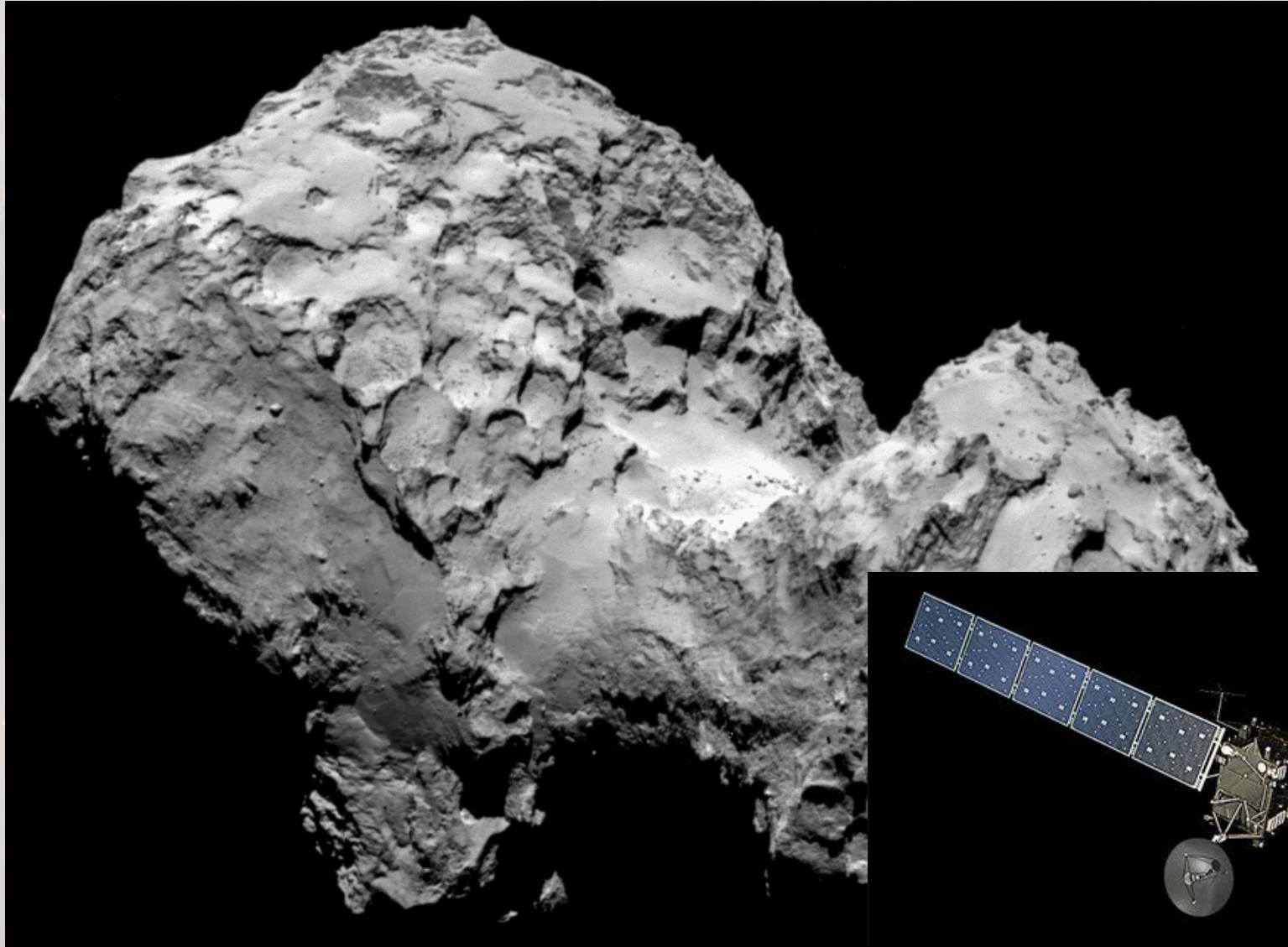


Rosetta - 67P



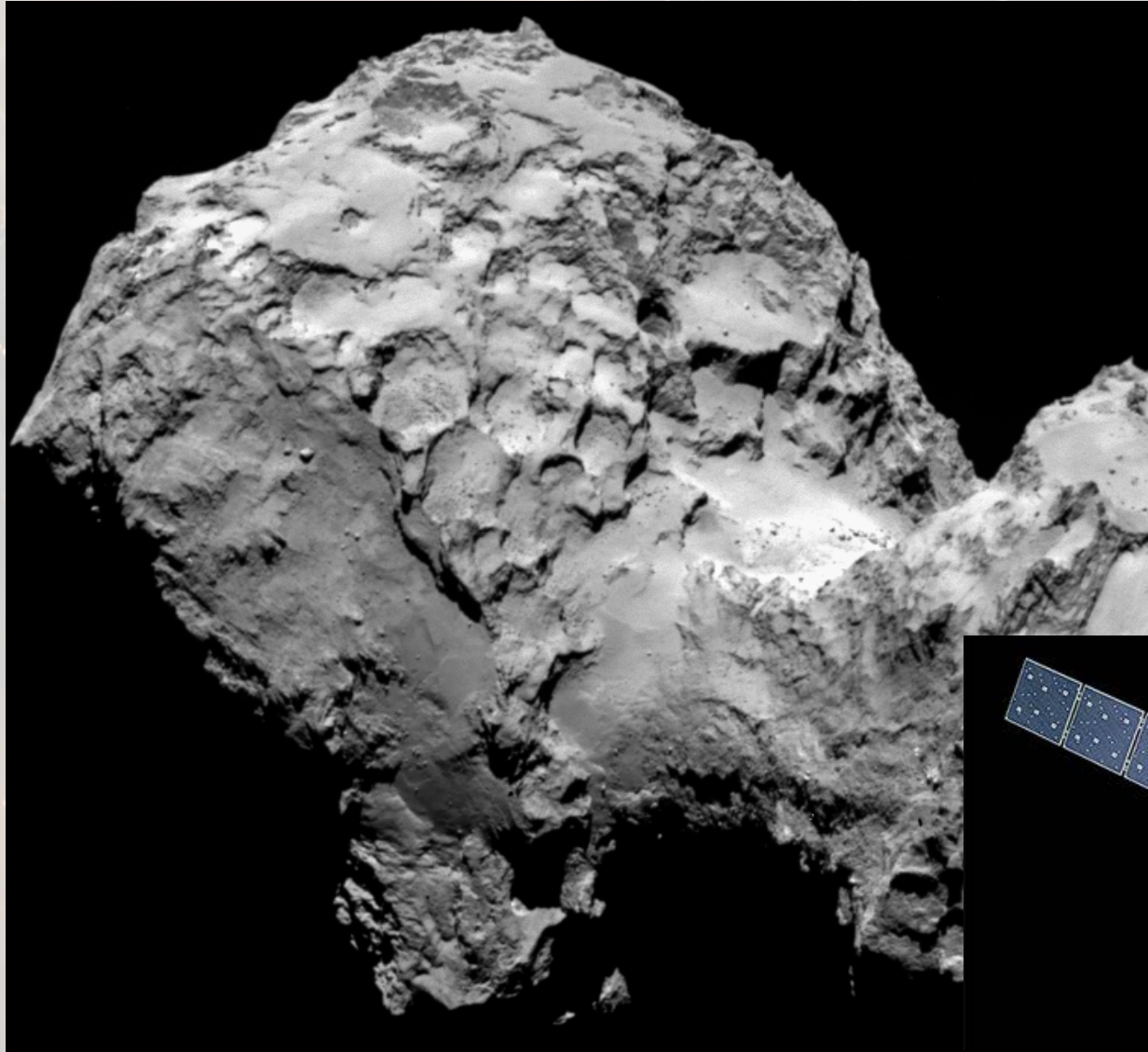
Credit: ESA – Rosetta Mission

Rosetta - 67P



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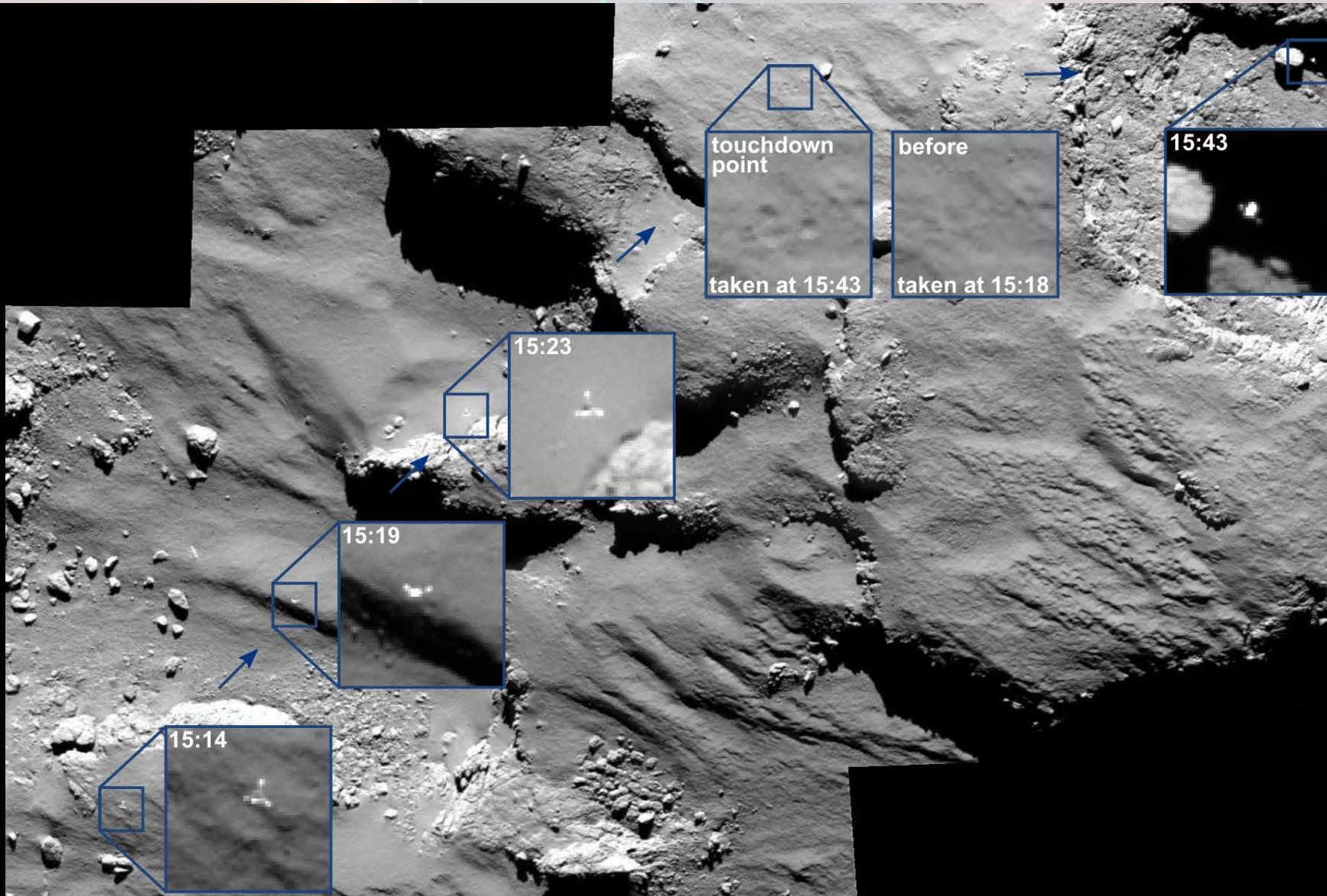


Credit: Nokia



Credit: ESA – Rosetta Mission

Rosetta - 67P



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Rosetta - 67P

$$\rho = M/V$$

$$V = \frac{4\pi R^3}{3}$$

$$\rho = 0.4 \text{ g/cm}^3$$

$$M_{67P} = 1 \times 10^{13} \text{ kg}$$

$$G = 6.67 \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2}$$

Rosetta - 67P

$$\rho = M/V$$

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$$\begin{aligned} \rho &= 0.4 \times 10^{-3} \text{ kg} / 1 \times 10^{-6} \text{ m}^3 \\ &= 0.4 \times 10^3 \text{ kg/m}^3 \end{aligned}$$

Rosetta - 67P

$$\rho = M/V$$

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$$M = \frac{4\pi R^3 \rho}{3}$$

$$1 \times 10^{13} = \frac{4\pi R^3 \rho}{3}$$

$$R = \sqrt[3]{\frac{3 * 1 \times 10^{13}}{4\pi * 0.4 \times 10^3}} = 2 \times 10^3 \text{ m}$$

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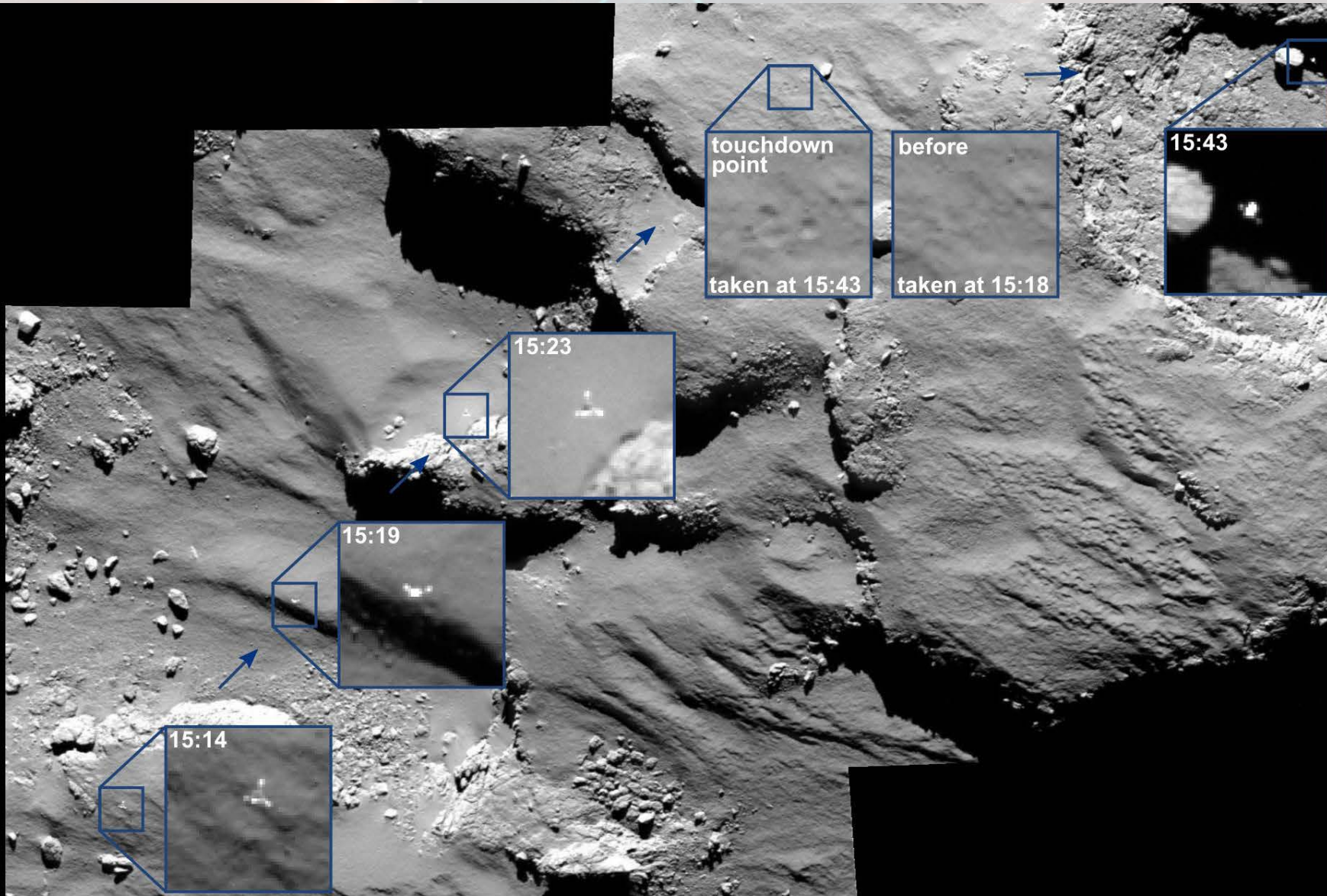
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$$g_c = 2 \times 10^{-4} \text{ ms}^{-2}$$

$$\frac{g_c}{g_E} = \frac{2 \times 10^{-4}}{9.8} = 2 \times 10^{-5}$$

Rosetta - 67P



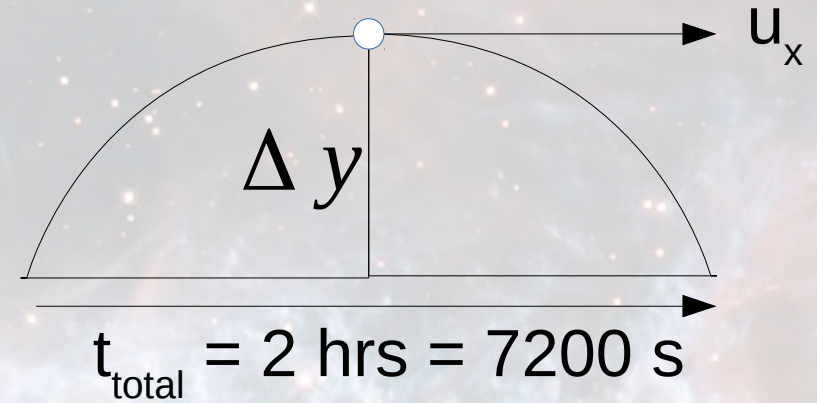
Credit: ESA – Rosetta Mission

Rosetta – 67P

Double checking with Galileo
and Projectile Motion

$$\Delta y = u_y t + \frac{1}{2} a t^2$$

$$\Delta y = 0 + \frac{1}{2} g_c t^2$$



$$\Delta y = 1 \text{ km}$$

$$u_y = 0 \text{ ms}^{-1}$$

$$t = 3600 \text{ s}$$

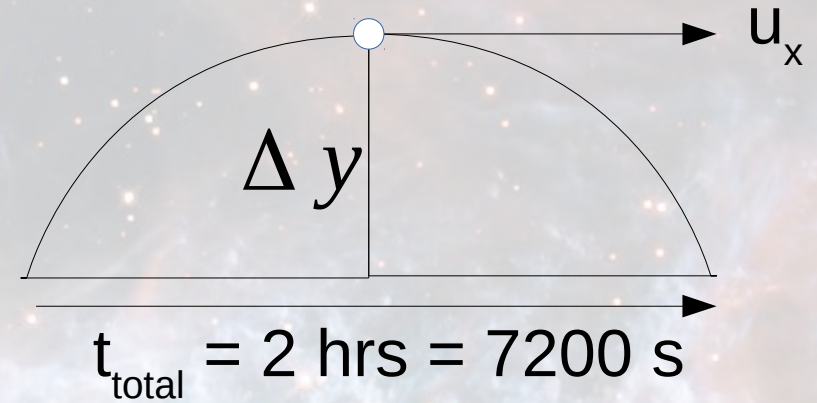
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Rosetta – 67P

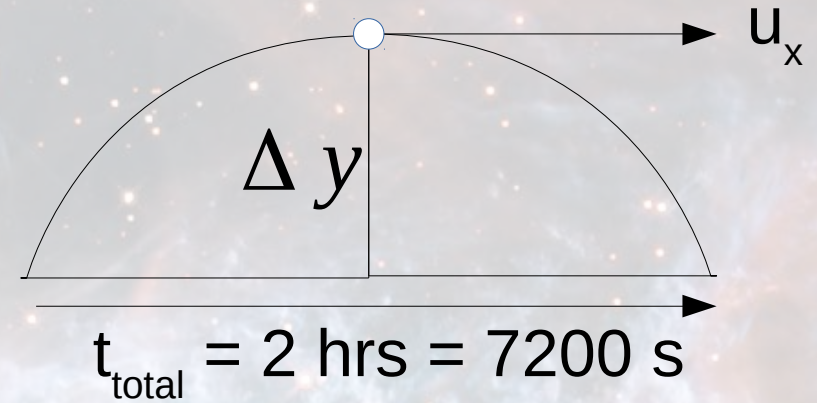
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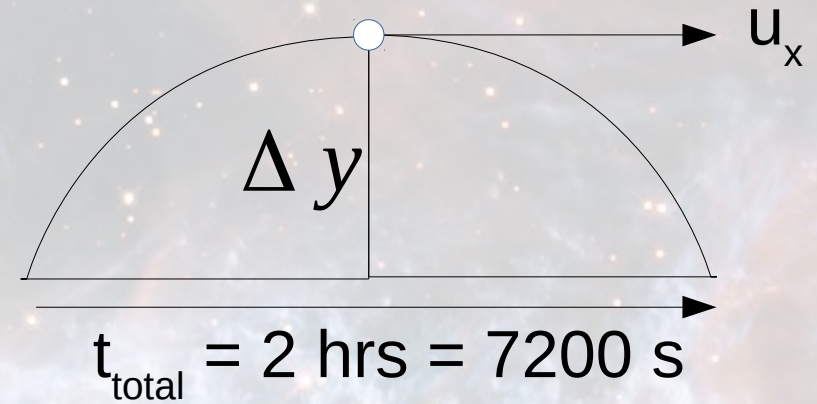
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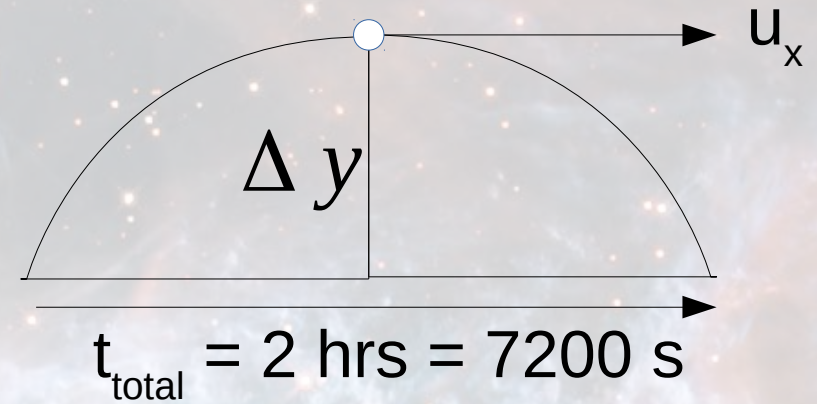
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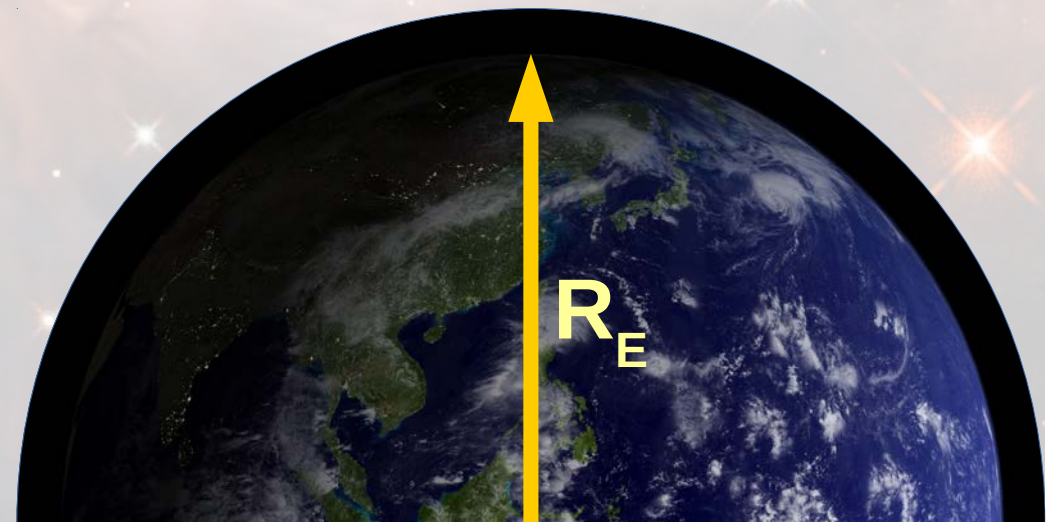
$$\frac{g_c}{g_E} = \frac{2 \times 10^{-4}}{9.8} = 2 \times 10^{-5}$$

Gravitational Potential Energy

$$W = Fd$$

$$W = F_g d$$

$$W = mgd$$



Gravitational Potential Energy

$$E_p = W_{d=\infty \rightarrow r}$$



Gravitational Potential Energy

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Gravitational Potential Energy

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$$W = F_g d$$

$$W = \frac{GM}{d^2} m d$$

$$E_p = 0 - \frac{GM}{r^2} m r$$



Gravitational Potential Energy

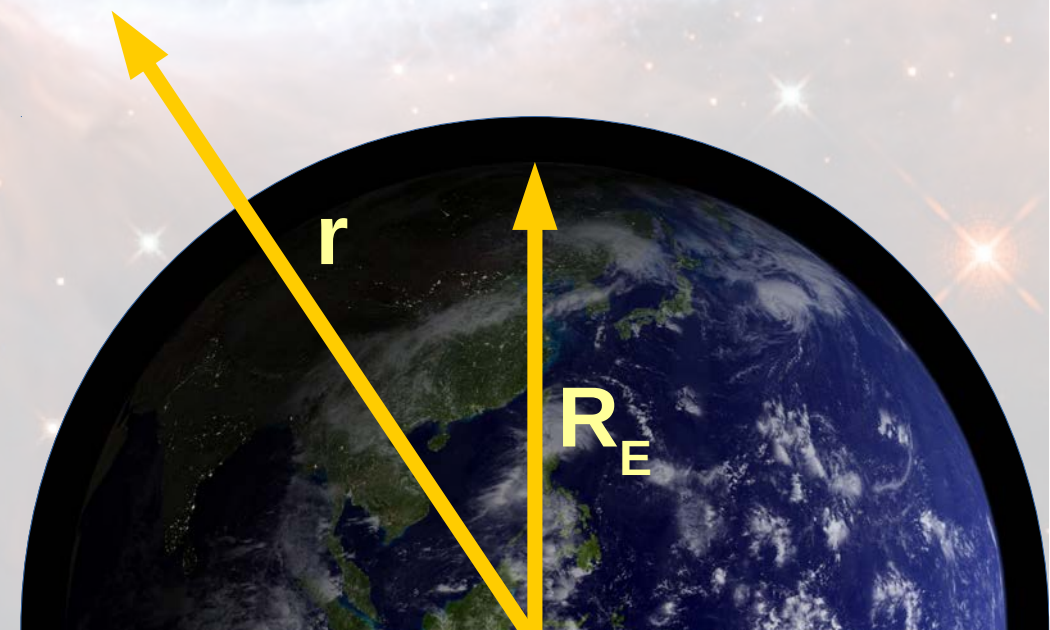
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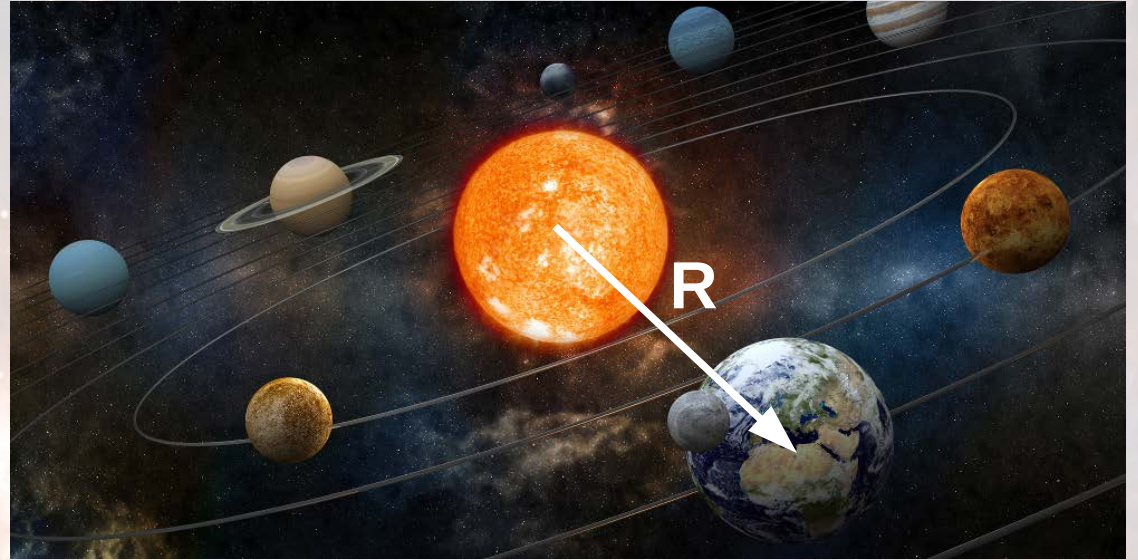
$$E_p = 0 - \frac{GM}{r^2} m r$$

$$E_p = -\frac{GM}{r} m$$



Gravitational Potential Energy

$$E_p = -\frac{GM_{\odot}}{R} M_E$$



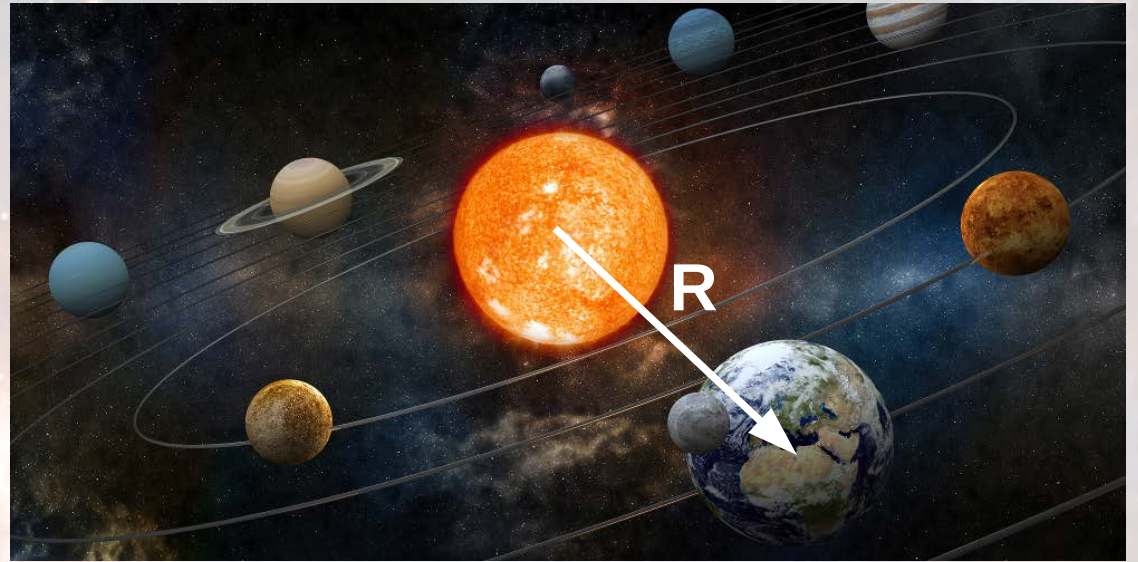
Gravitational Potential Energy

$$E_p = -\frac{GM_{\odot}}{R} M_E$$

$$M_{\odot} = 1.989 \times 10^{30} \text{ kg}$$

$$M_E = 5.97 \times 10^{24} \text{ kg}$$

$$R = 1.5 \times 10^8 \text{ km}$$



Gravitational Potential Energy

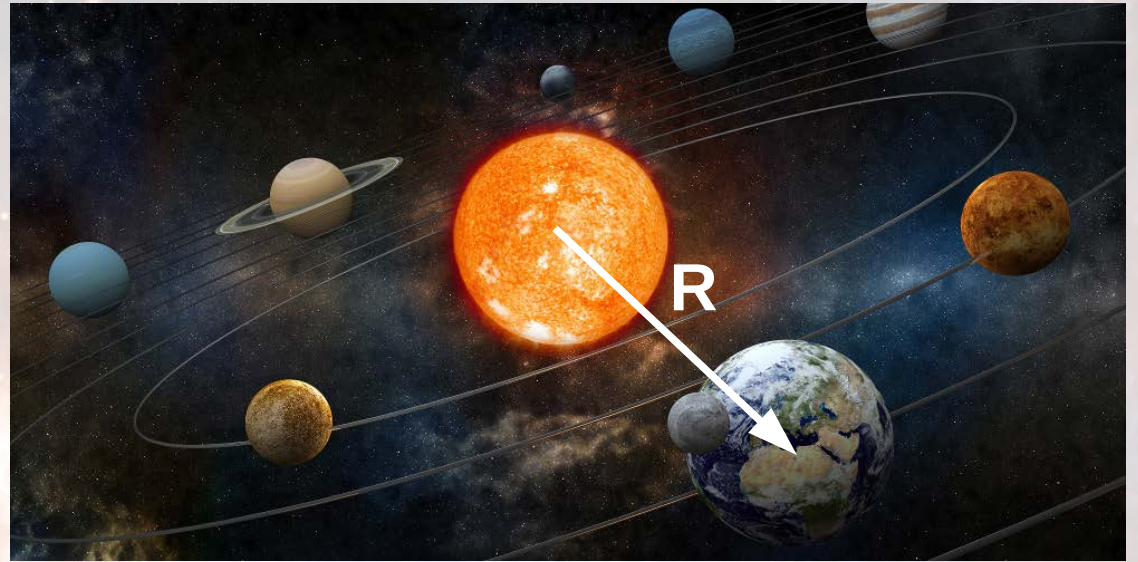
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Gravitational Potential Energy

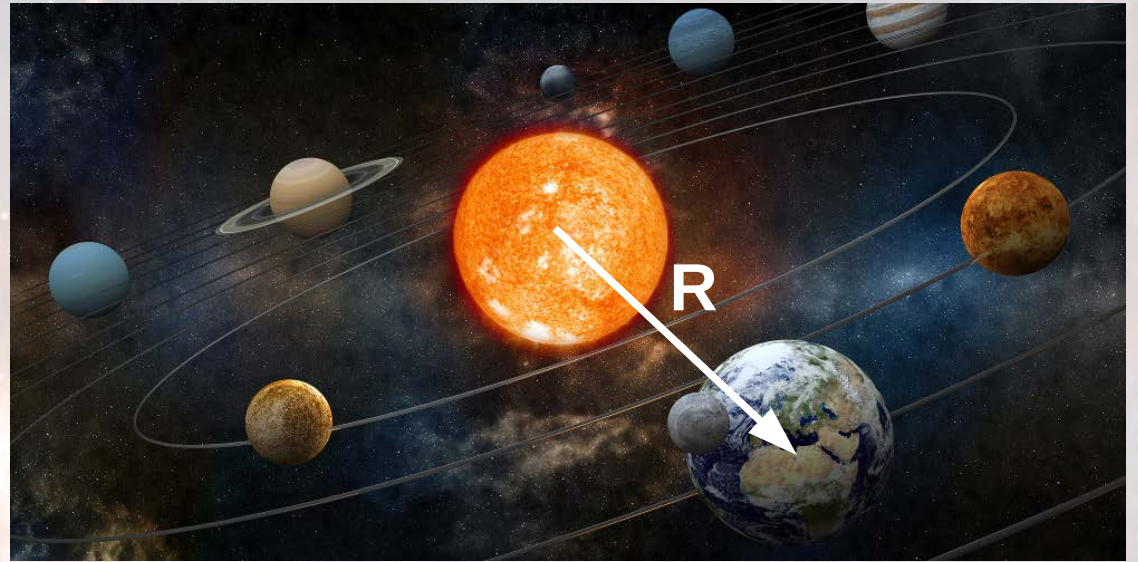
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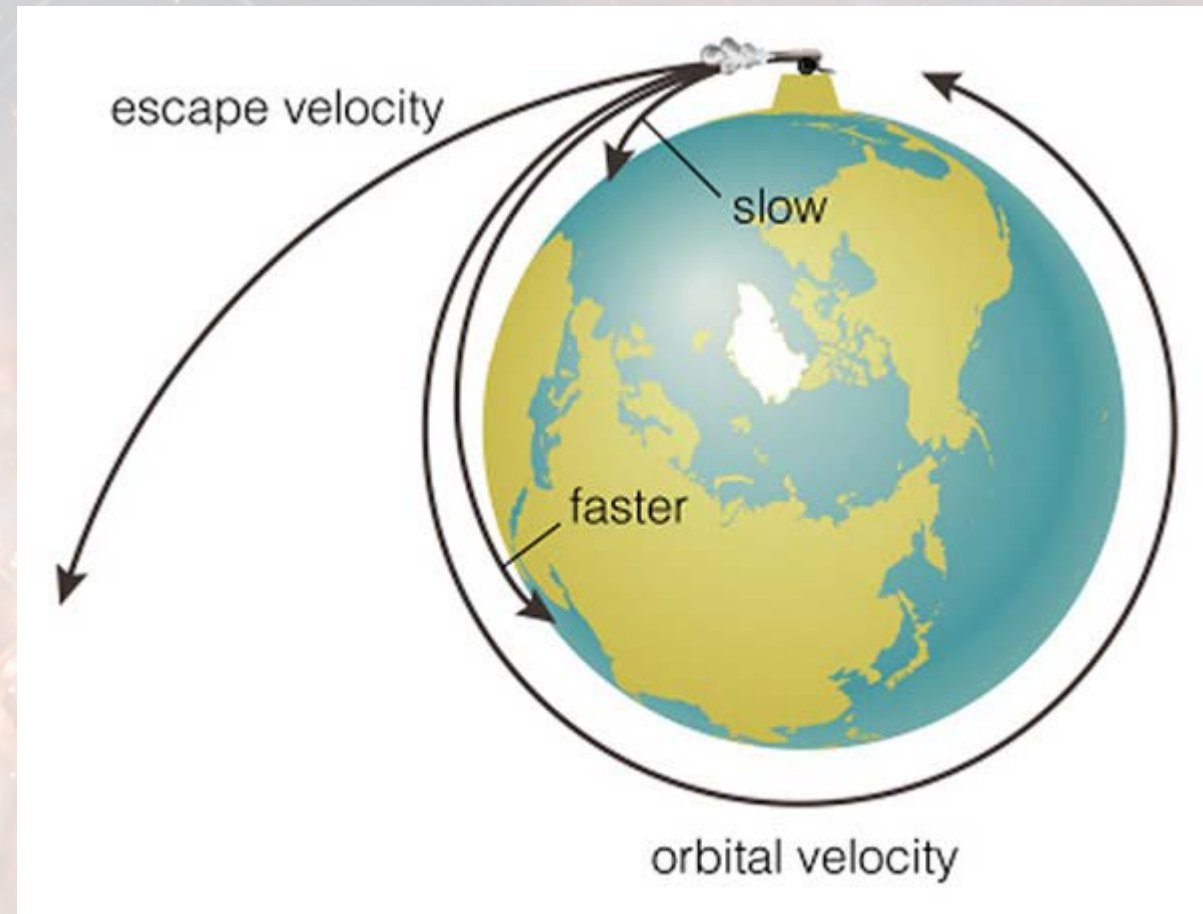


Equivalent to setting off every nuke on Earth every second for 250 000 years or the energy released from the surface of the Sun in 160 days.

Rocket Science!

To reach orbit:

$$KE = E_p$$



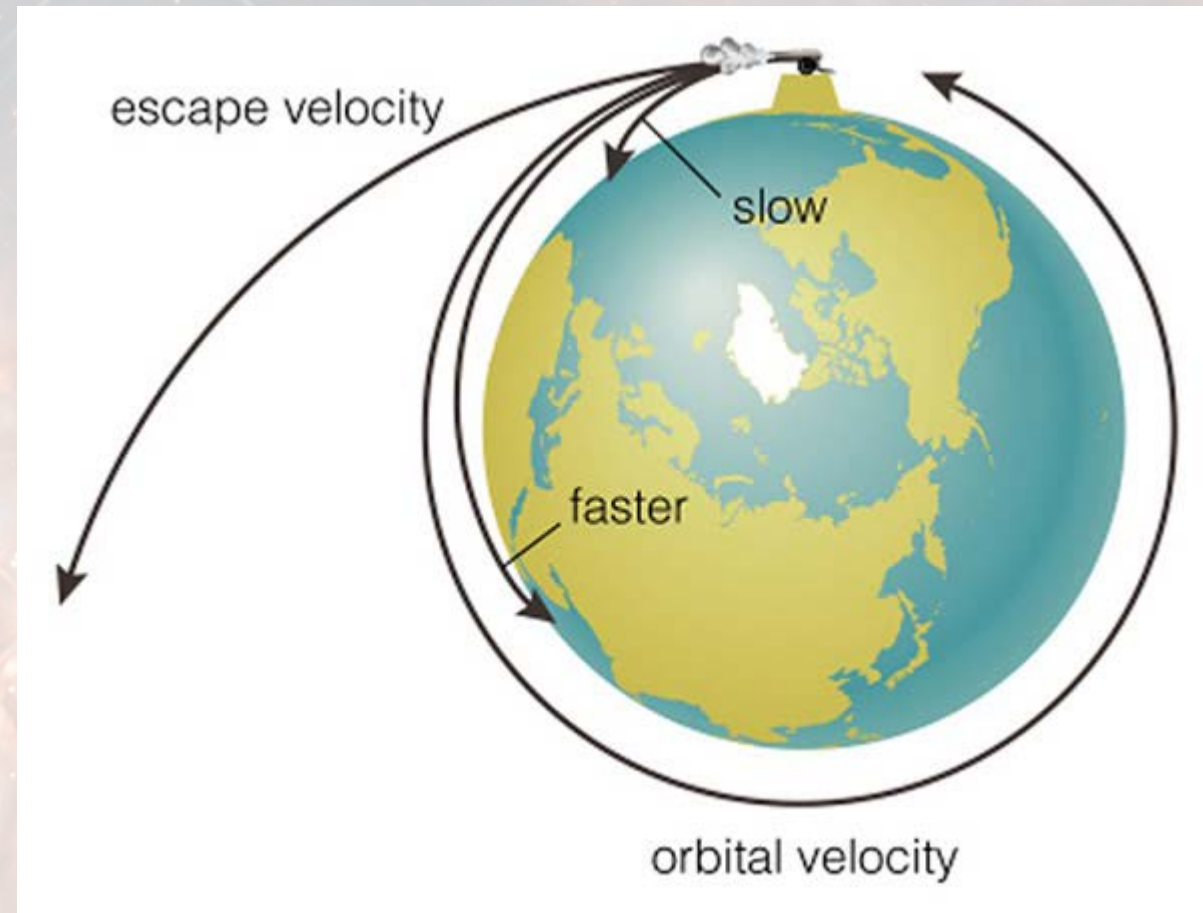
Credit: ustudy.in

Rocket Science!

To reach orbit:

$$KE = E_p$$

$$\frac{mv^2}{2} = \frac{GMm}{r}$$



Credit: ustudy.in

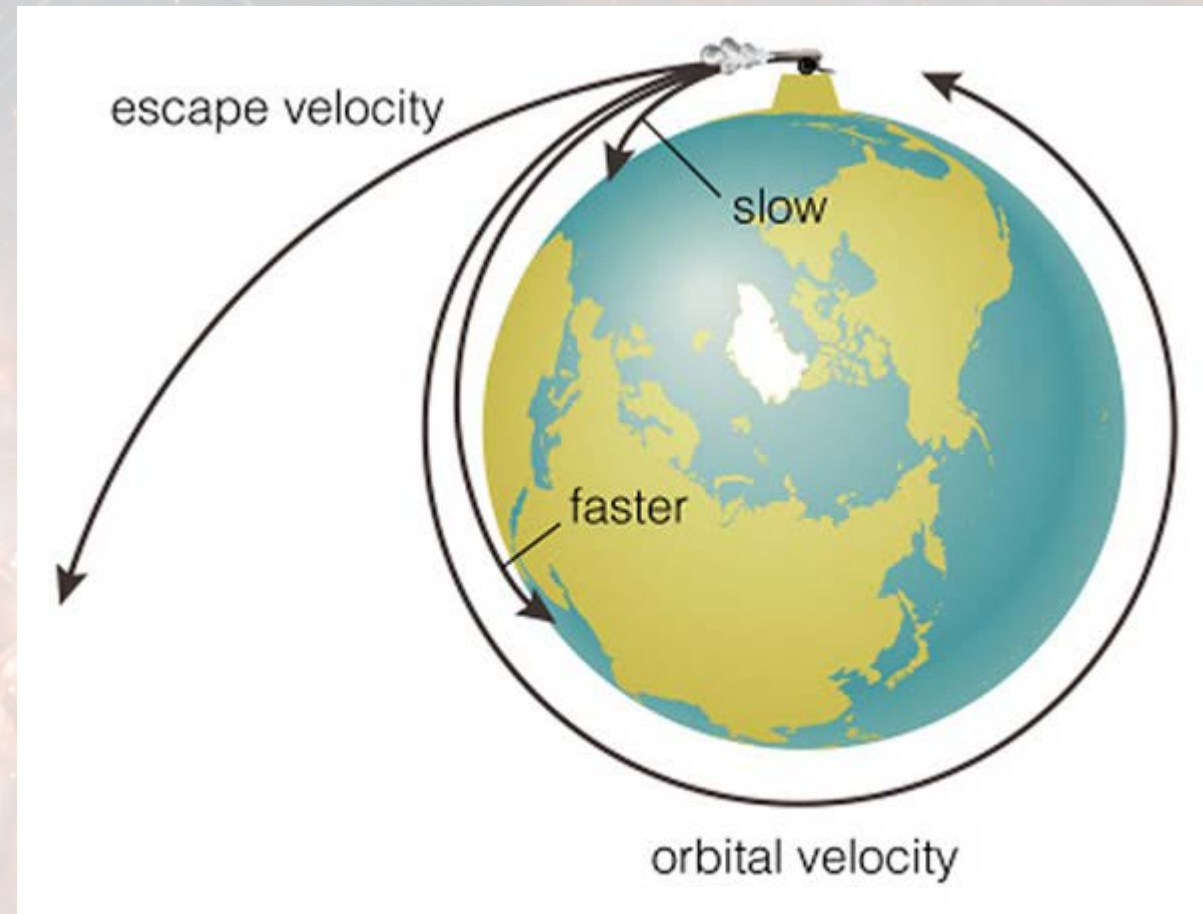
Rocket Science!

To reach orbit:

$$KE = E_p$$

$$\frac{mv^2}{2} = \frac{GMm}{r}$$

$$v_{orb} = \sqrt{\frac{2GM}{r}}$$



Credit: ustudy.in

Rocket Science!

To reach escape velocity:

$$KE \geq E_p$$



Credit: hdwallpapers.in

Rocket Science!

To reach escape velocity:

$$KE \geq E_p$$

$$\frac{mv^2}{2} \geq \frac{GMm}{r}$$



Credit: hdwallpapers.in

Rocket Science!

To reach escape velocity:

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$$v_{esc} \geq \sqrt{\frac{2GM}{r}}$$



Credit: hdwallpapers.in

Rocket Science!

To reach escape velocity:

$$KE \geq E_p$$

$$\frac{mv^2}{2} \geq \frac{GMm}{r}$$

$$v_{esc} \geq \sqrt{\frac{2GM}{r}}$$

$$v_{esc} \geq \sqrt{\frac{2GM_{\odot}}{R_{1AU}}} \geq 42000 \text{ ms}^{-1}$$



Credit: hdwallpapers.in

