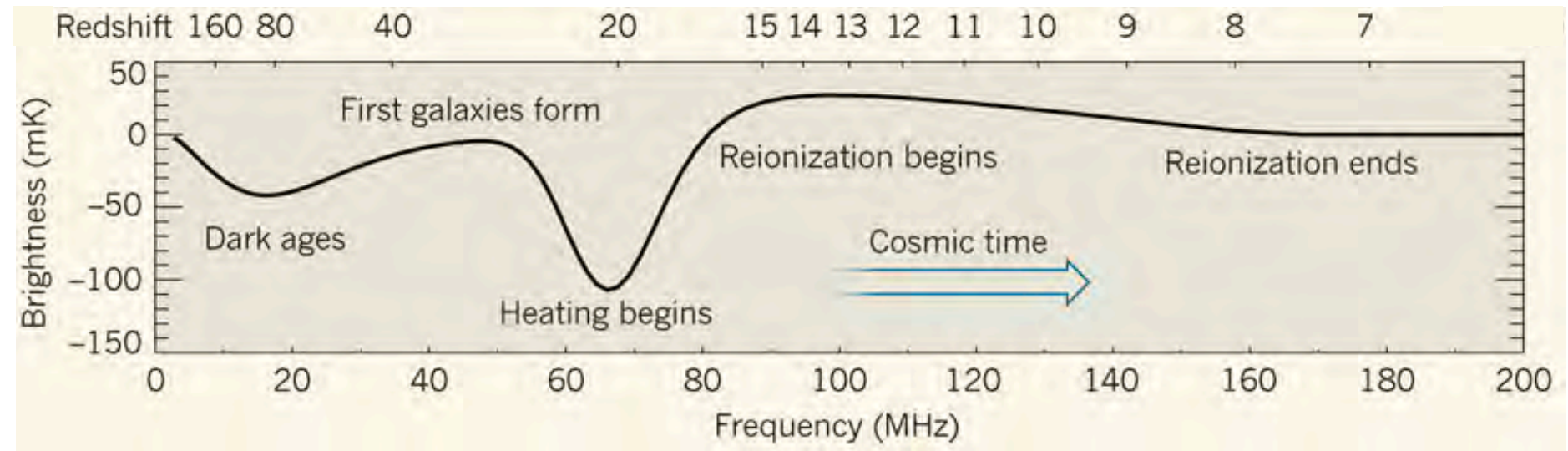


The Current State of Global Signal Theory



Pritchard & Loeb, *Nature*, 468, 722 (2010)

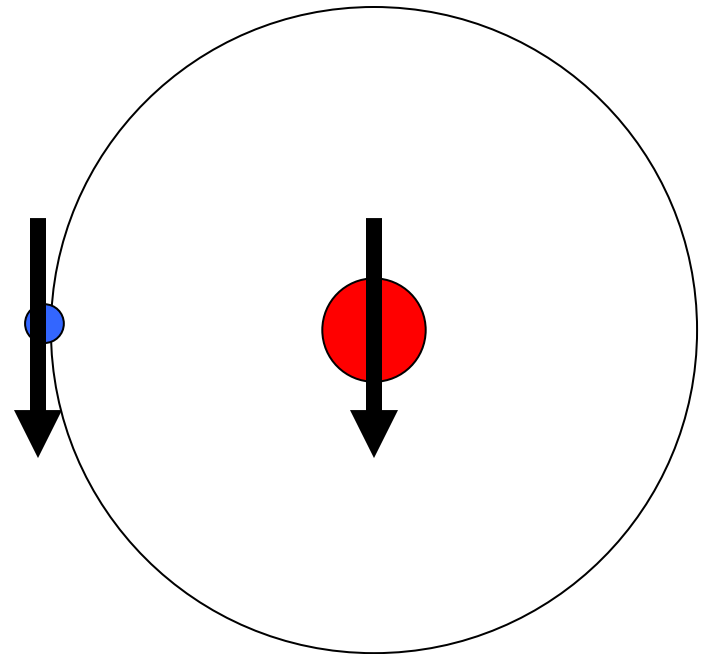
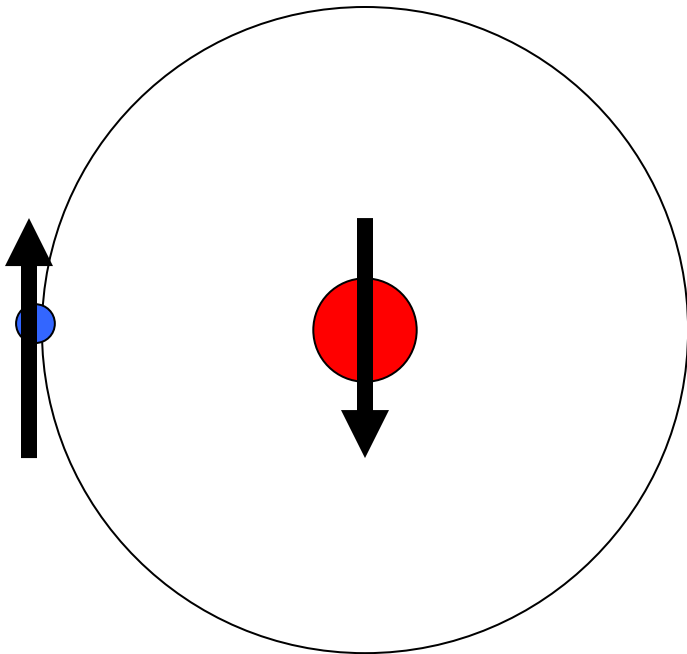
CAASTRO, 19th November 2012

Adrian Liu

Recent theoretical developments do
not change the basic picture

$$\frac{n_1}{n_0} = \frac{g_1}{g_0} e^{-E_{10}/k_B T_S}$$

Spin Temp



Neutral hydrogen
fraction

CMB temperature

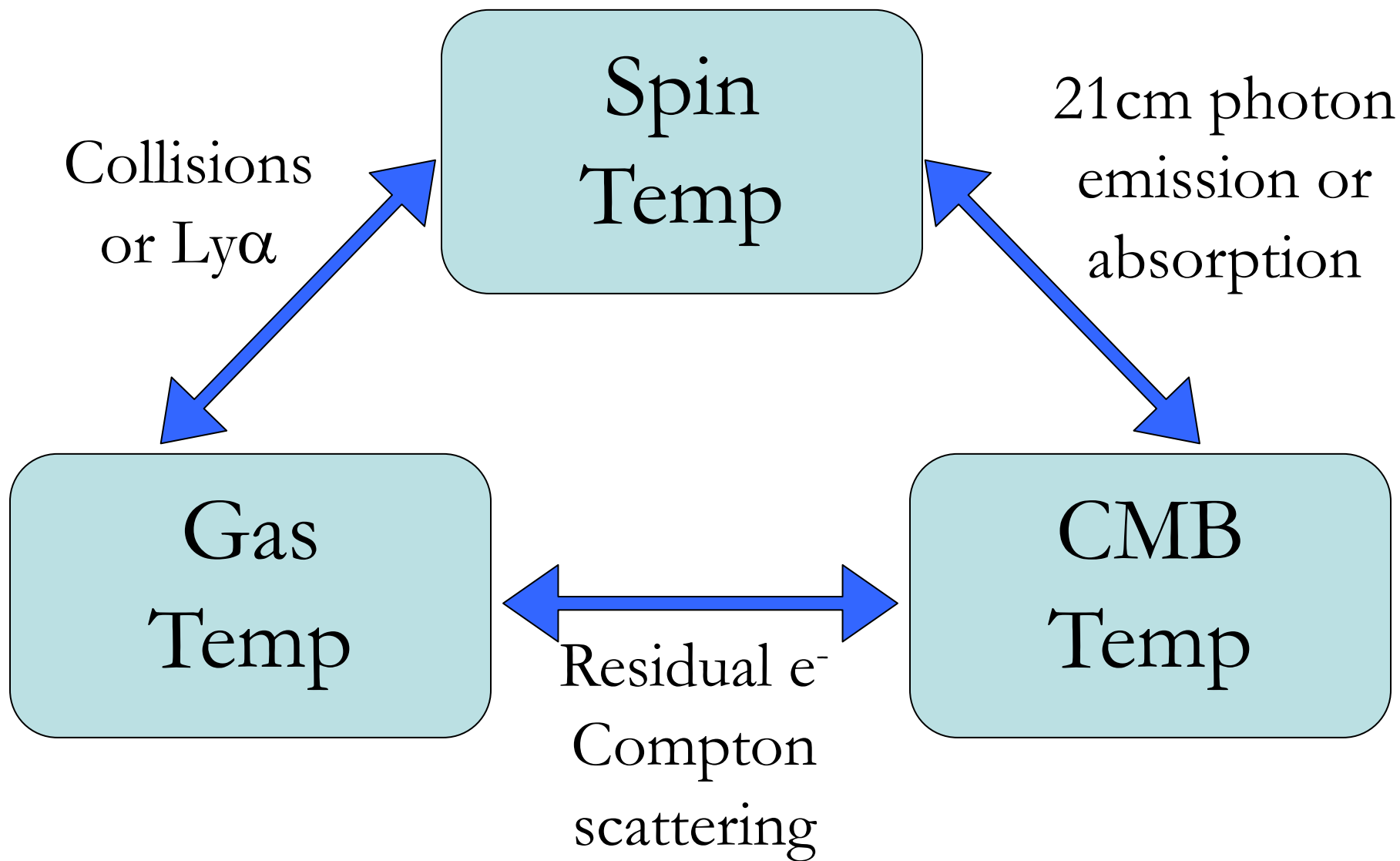
$$T_b = 27x_H \left(\frac{T_S - T_\gamma}{T_S} \right) \left(\frac{1+z}{10} \right)^{1/2} \text{mK}$$

21cm
brightness
temperature

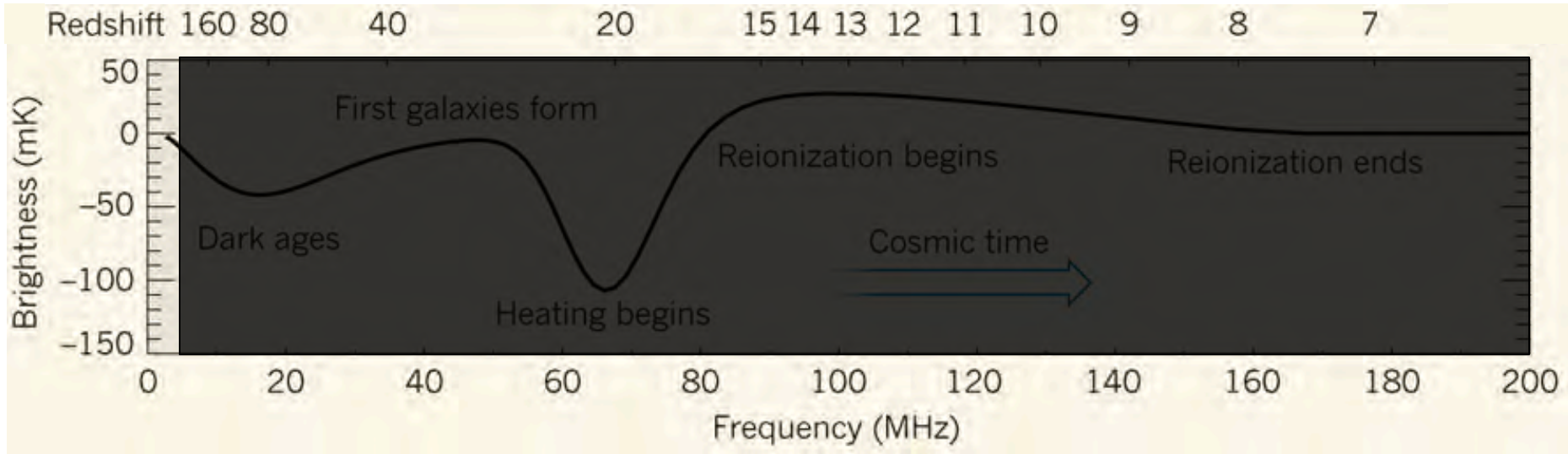
Spin
temperature

$$T_b = 27x_H \left(\frac{T_S - T_\gamma}{T_S} \right) \left(\frac{1+z}{10} \right)^{1/2} \text{mK}$$

Did we get
this right?



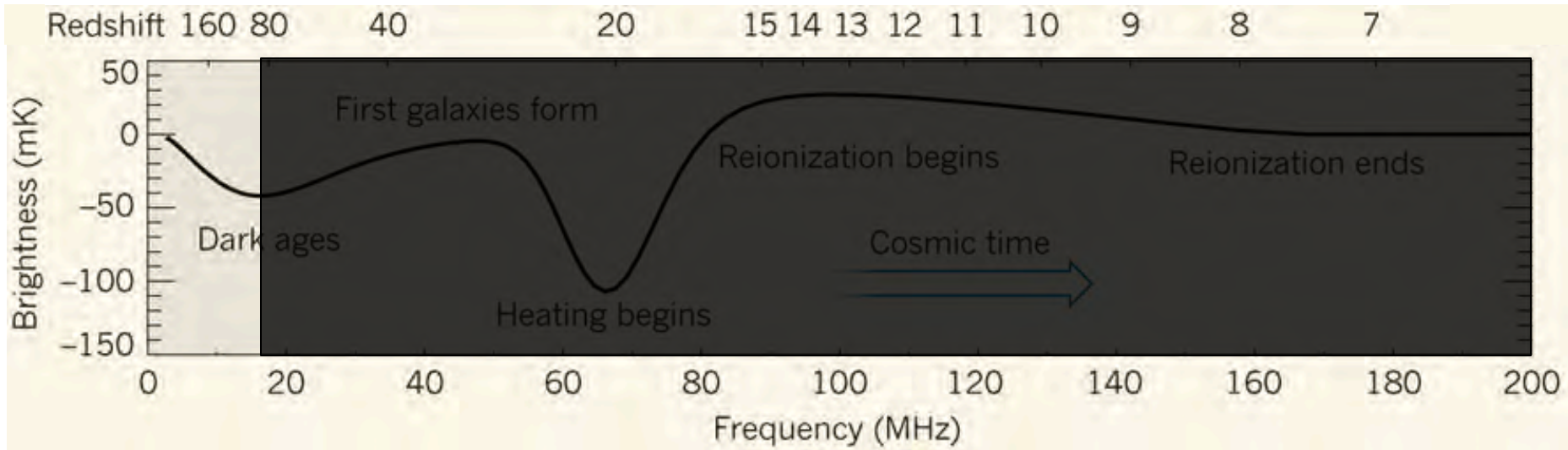
Phase I: No Signal



- Residual electrons drive T_{gas} to T_{CMB} via Compton scattering
- Collisions drive T_{gas} to T_{spin}

NO SIGNAL

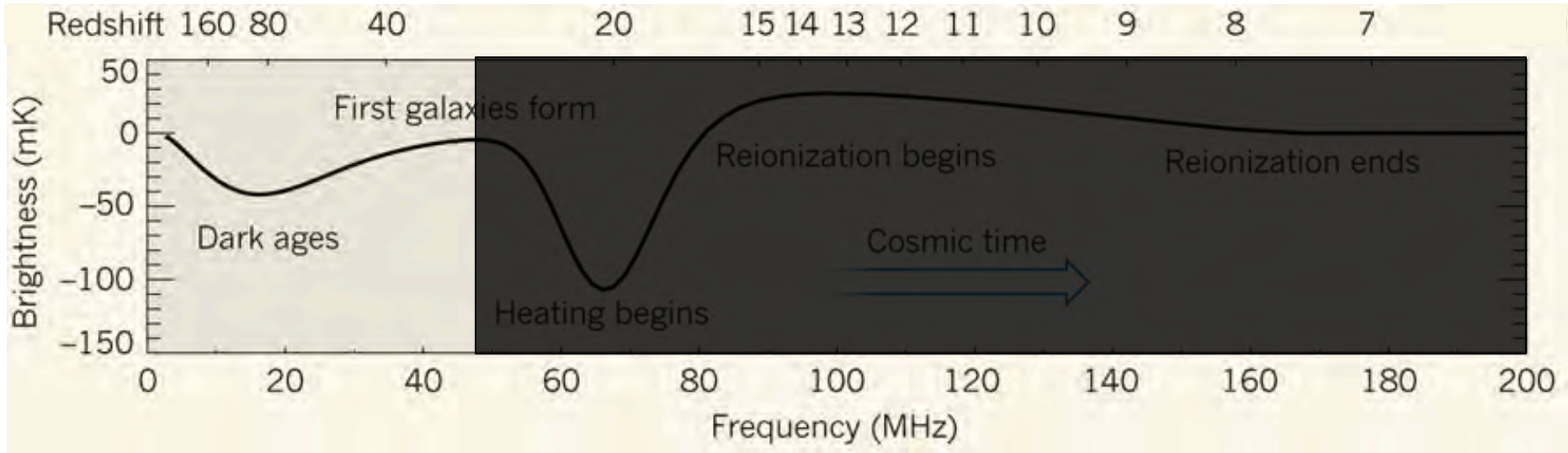
Phase II: Absorption Signal



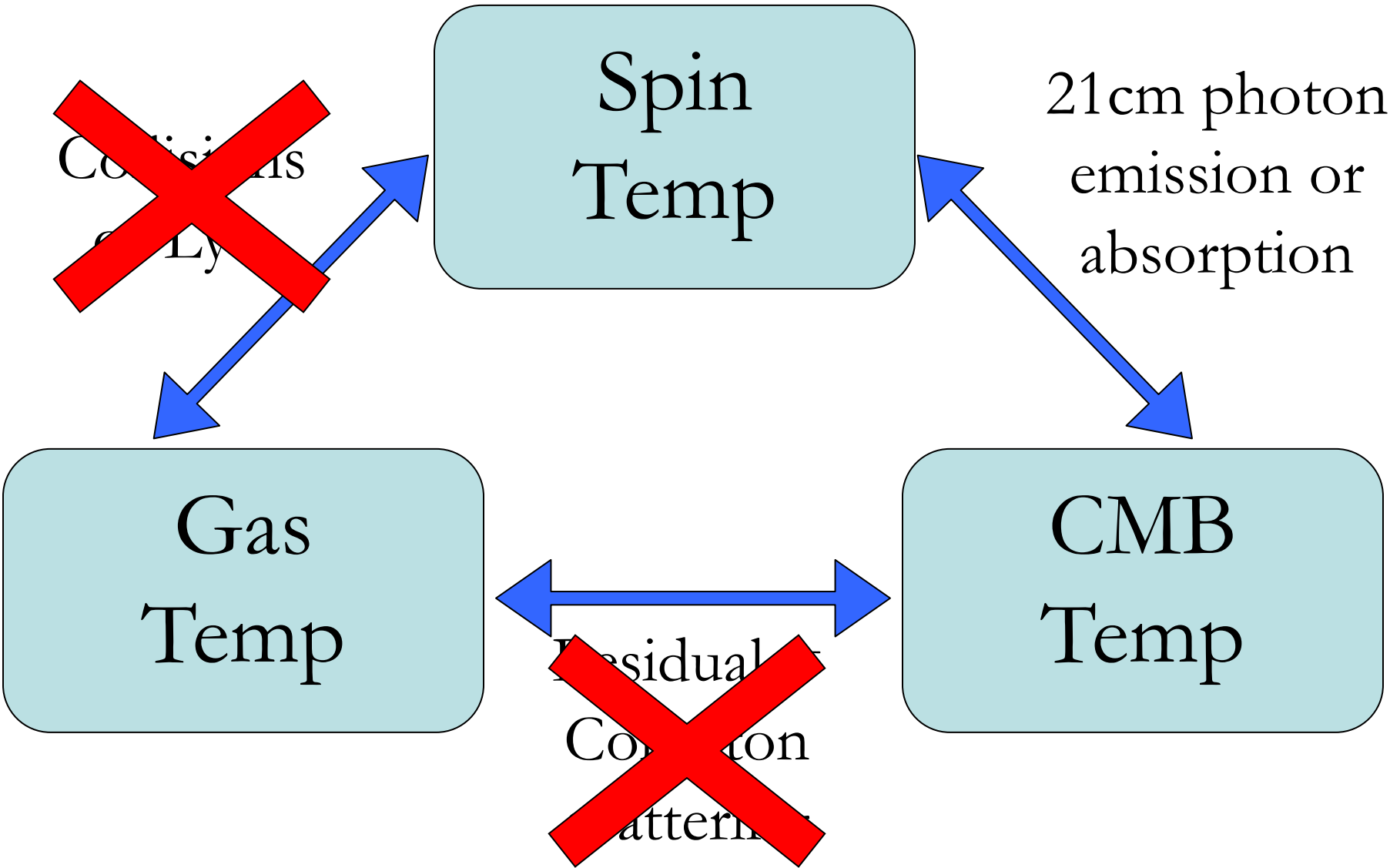
- Collisions maintain T_{gas} and T_{spin} equilibrium
- CMB has become too dilute for T_{CMB} and T_{gas} coupling
- $T_{\text{spin}} = T_{\text{gas}} \sim a^{-2}$ while $T_{\text{CMB}} \sim a^{-1}$

ABSORPTION

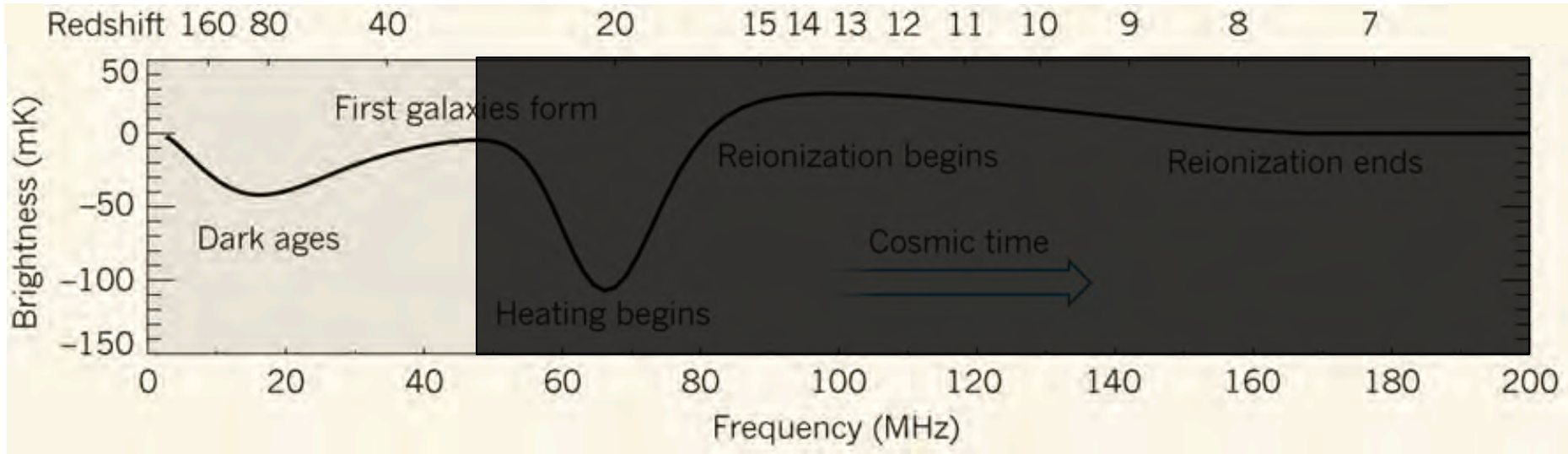
Phase III: Back to nothing



- Gas becomes too dilute for collisions to keep T_{gas} and T_{spin} in equilibrium



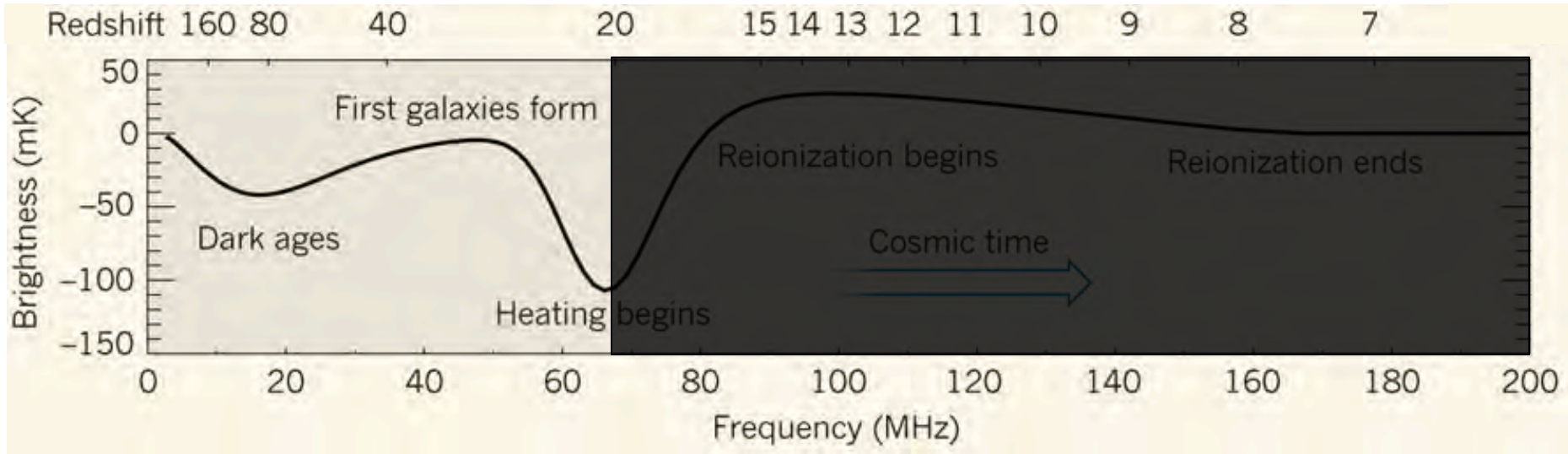
Phase III: Back to nothing



- Gas becomes too dilute for collisions to keep T_{gas} and T_{spin} in equilibrium
- Absorp./emission of 21cm photons makes $T_{\text{spin}} = T_{\text{CMB}}$

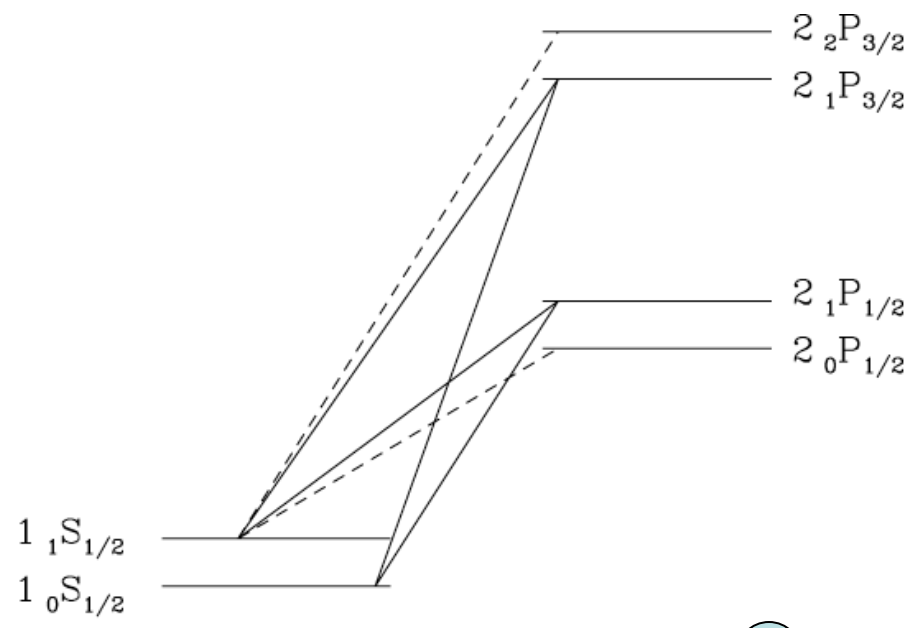
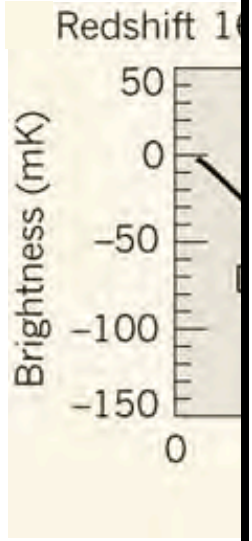
NO SIGNAL

Phase IV: First stars/galaxies form



- First stars create Ly α photons, causes coupling between T_{gas} and T_{spin} via Wouthuysen-Field effect

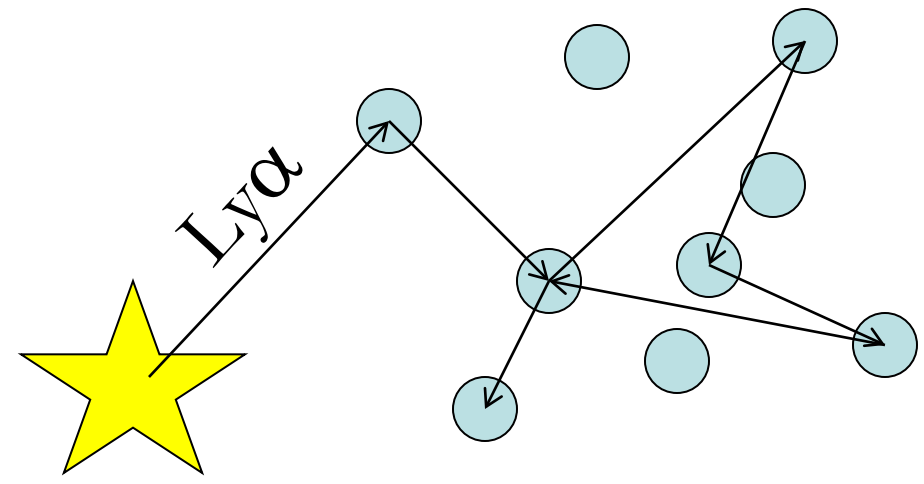
Ph



m

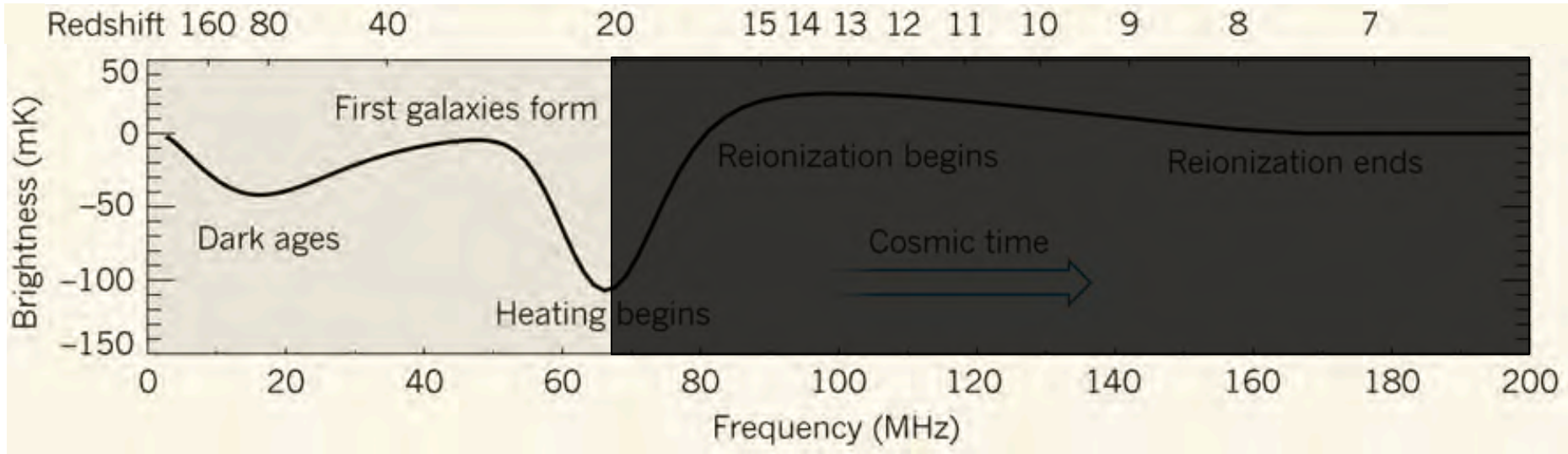


- First T_{gas}



een

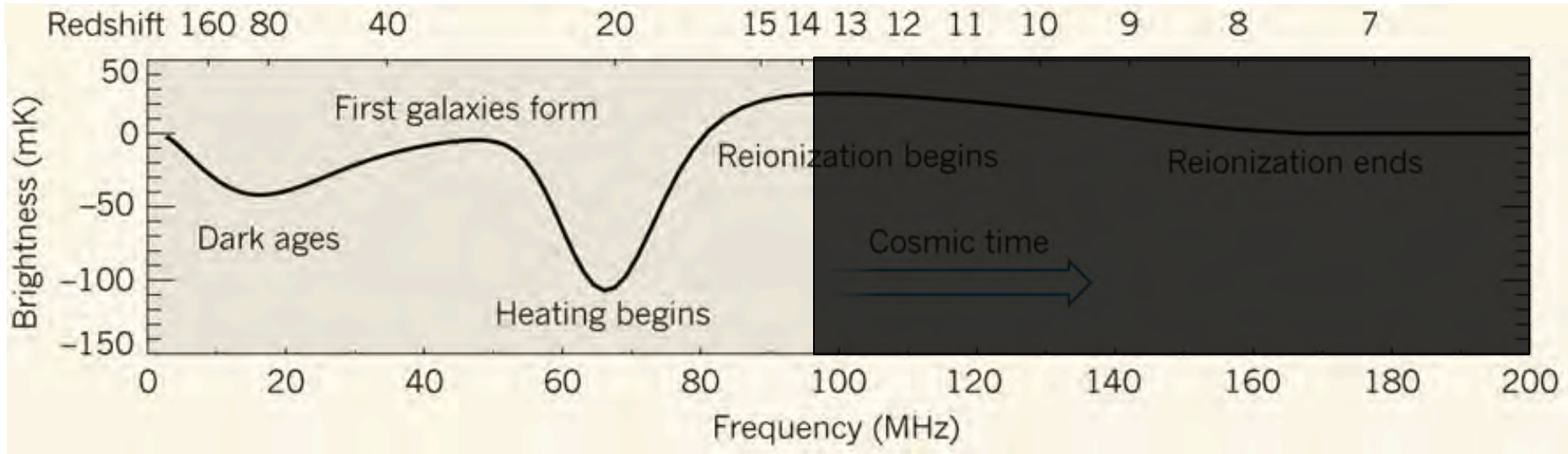
Phase IV: First stars/galaxies form



- First stars create Ly α photons, causes coupling between T_{gas} and T_{spin} via Wouthuysen-Field effect
- $T_{\text{spin}} = T_{\text{gas}} < T_{\text{CMB}}$, so...

ABSORPTION

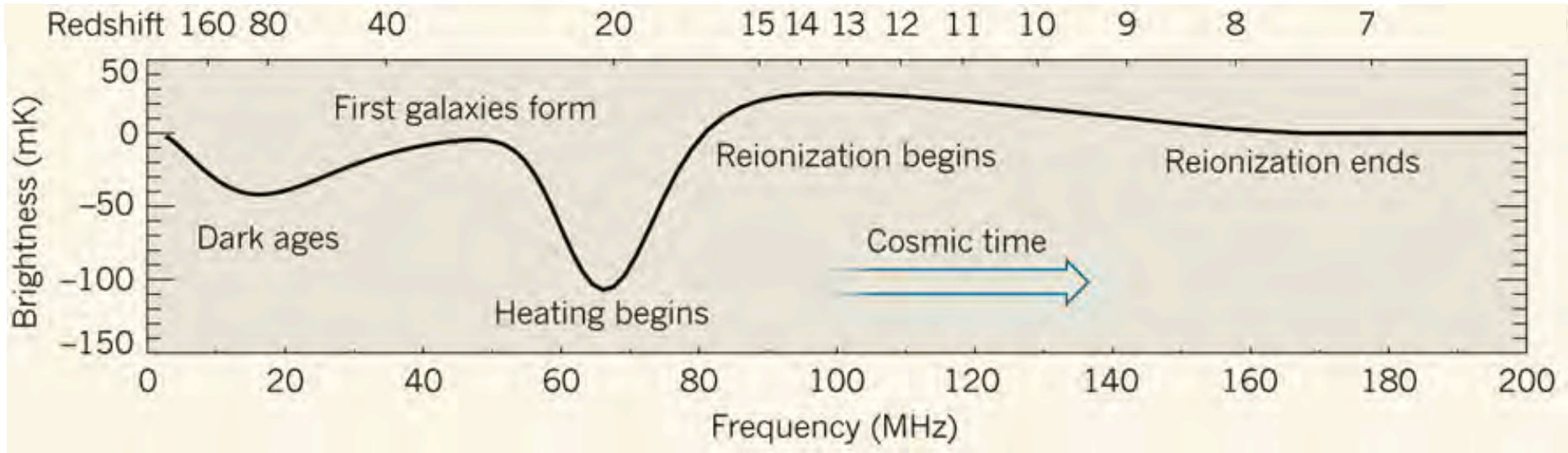
Phase V: X-ray reheating



- X-rays from early AGNs heat the IGM, raising T_{gas}
- Electrons are now going insane in atoms, giving off all sorts of photons causing $T_{\text{spin}} = T_{\text{gas}} > T_{\text{CMB}}$

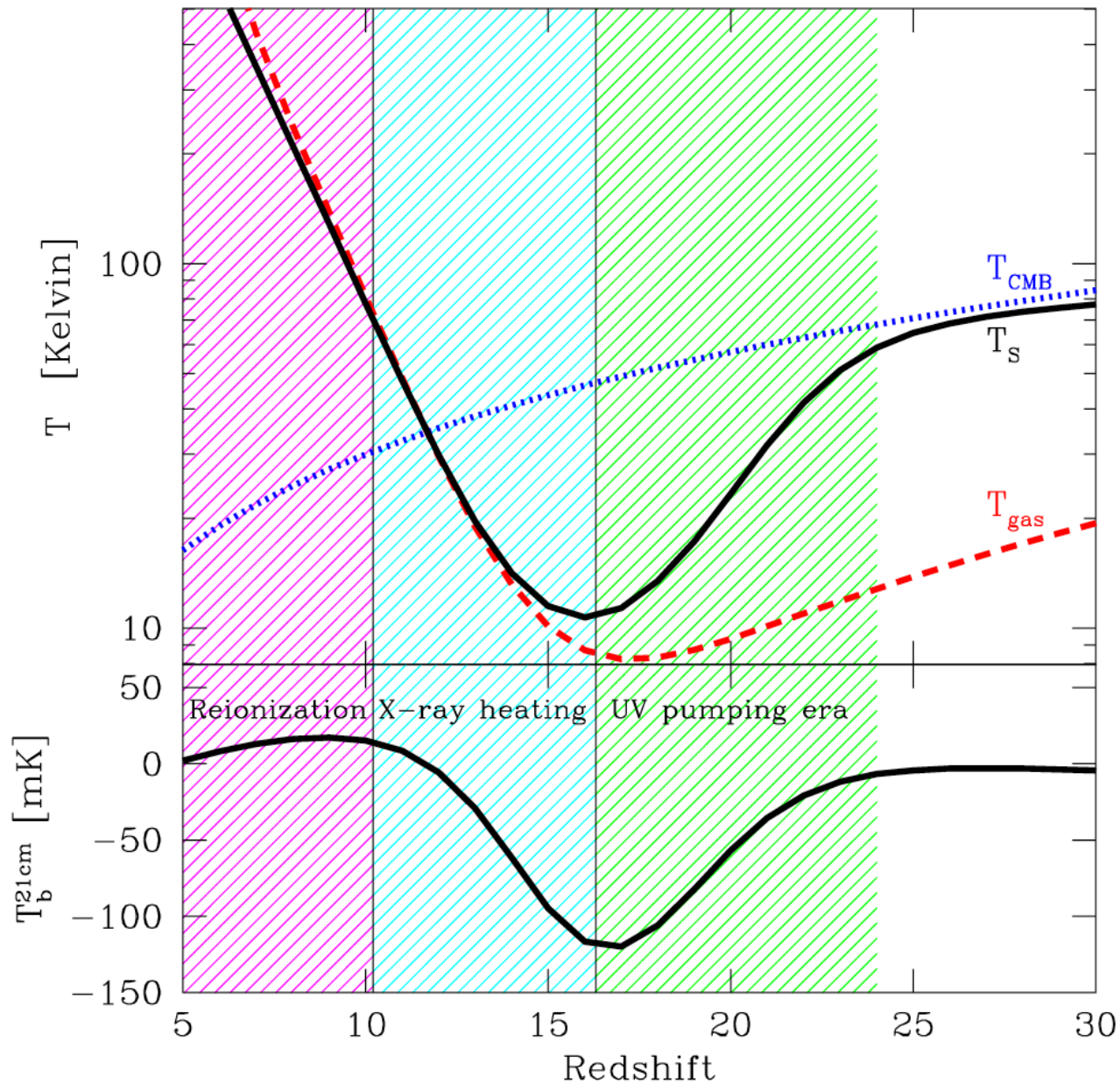
EMISSION

Phase VI: Reionization



- Eventually, reionization rids us of all neutral hydrogen atoms.

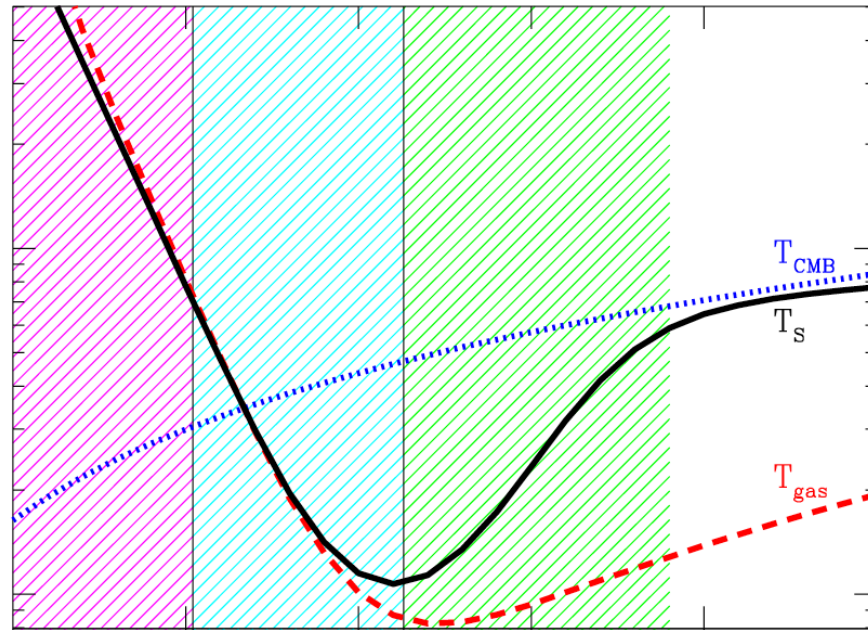
NO MORE SIGNAL



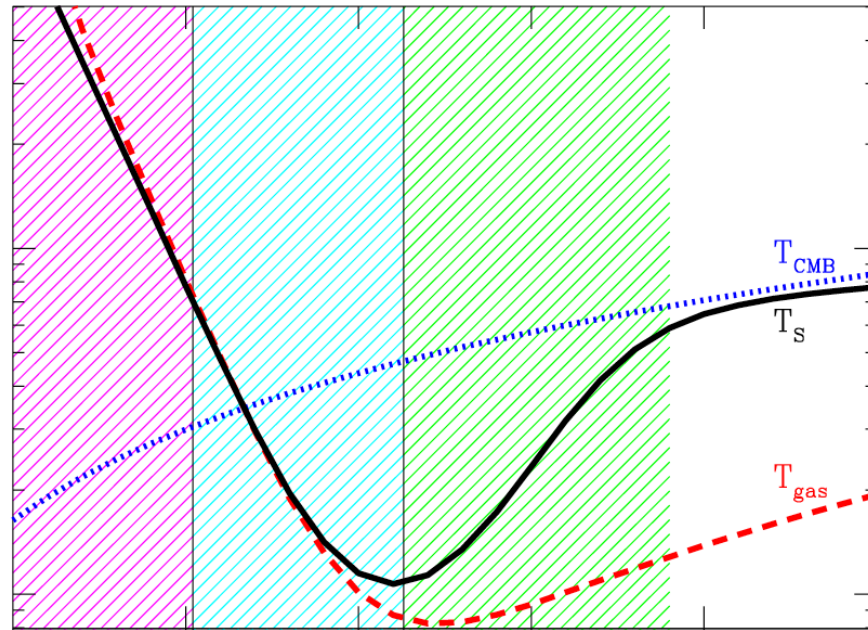
McQuinn & O'Leary (2012)

Nightmare scenario:
 might the absorption feature vanish with better modeling?

What might go wrong?

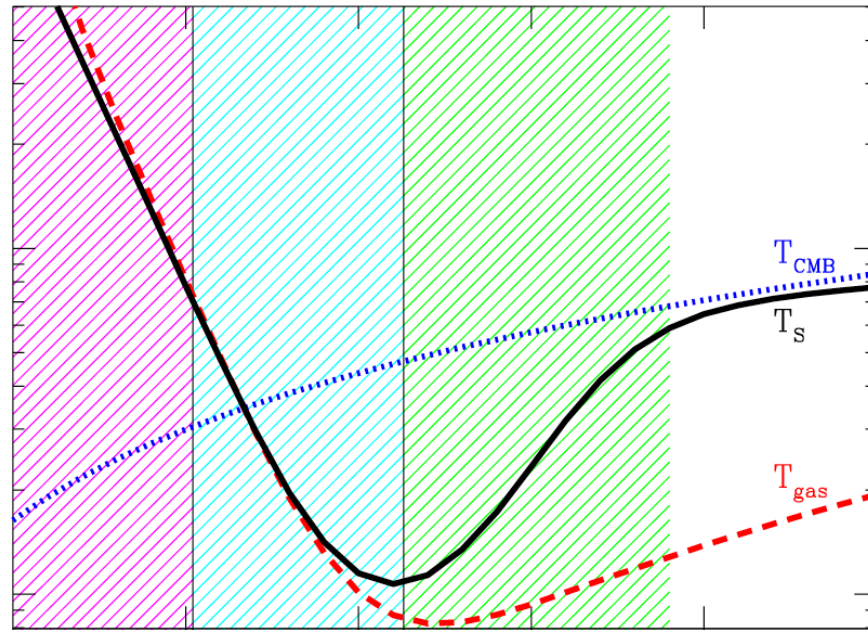


What might go wrong?



1. IGM physics may cause T_{gas} to deviate from adiabatic cooling.
2. X-ray heating may prematurely raise T_{spin} .

What might go wrong?



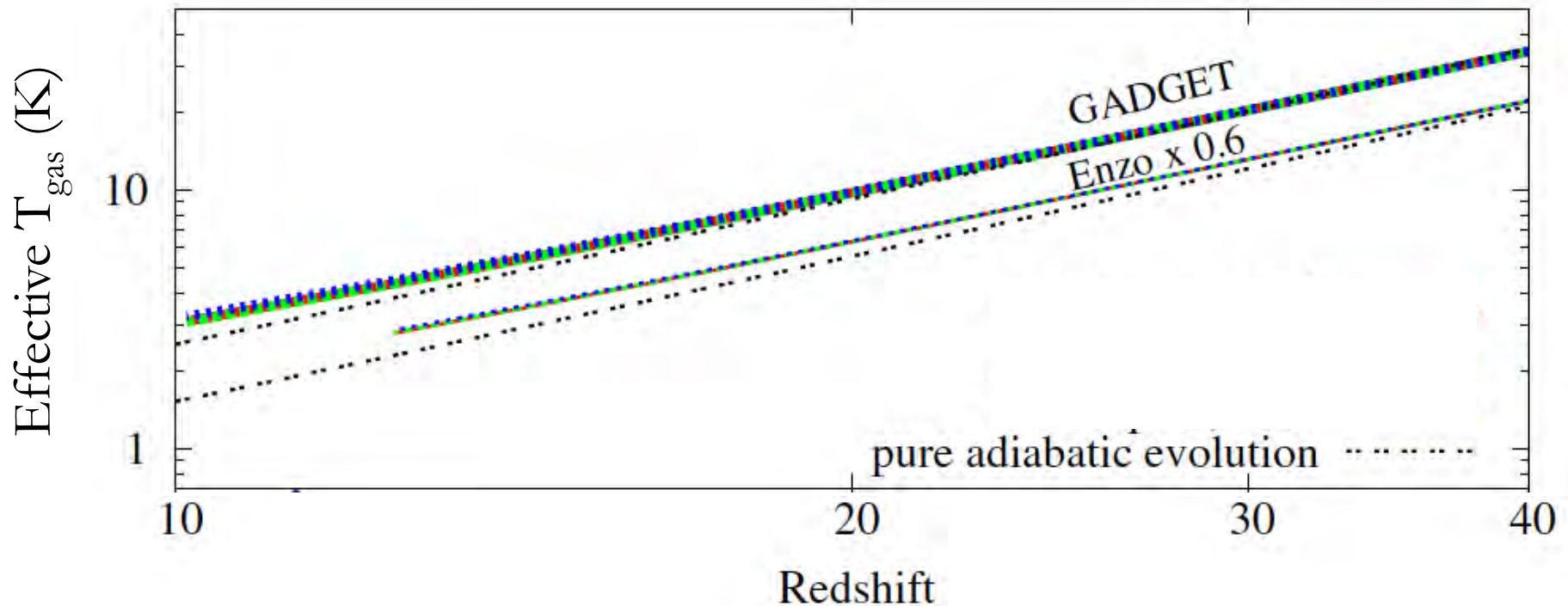
1. IGM physics may cause T_{gas} to deviate from adiabatic cooling.
2. X-ray heating may prematurely raise T_{spin} .

Supersonic flows could shock-heat
the IGM at high redshifts

Supersonic flows could shock-heat the IGM at high redshifts

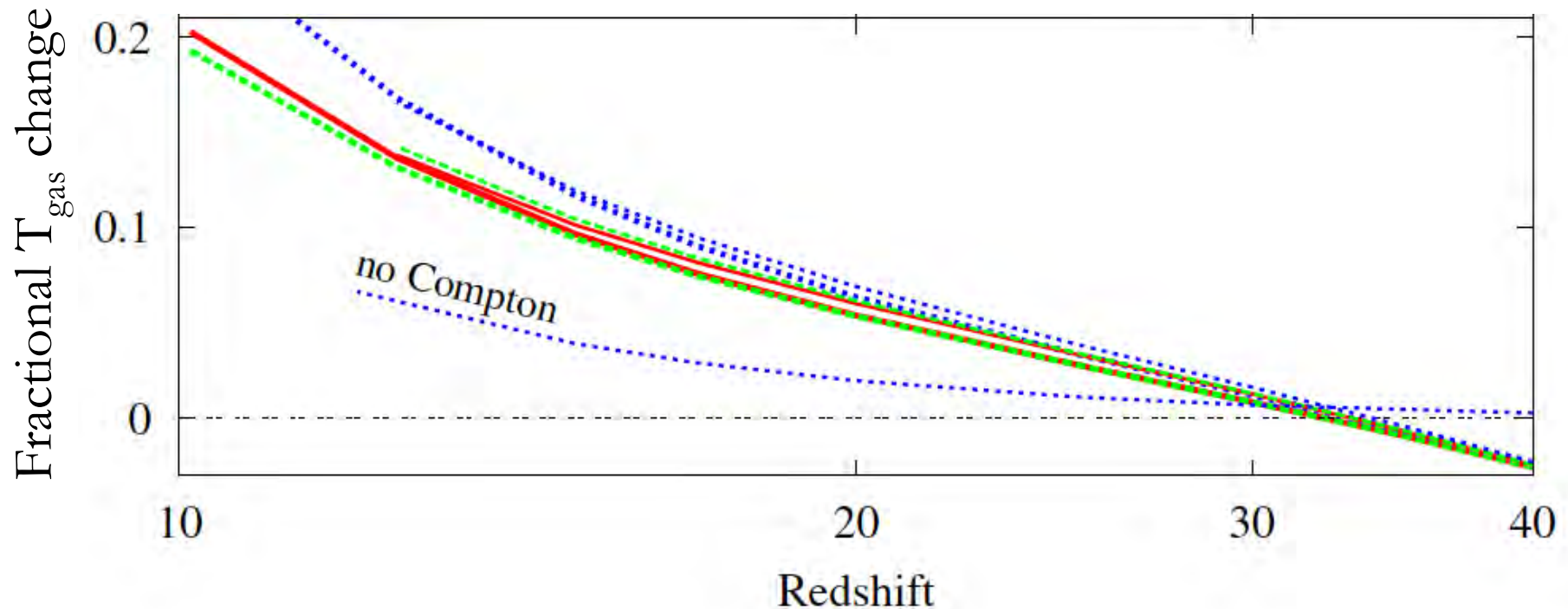
- Tseliakovich & Hirata (2010):
 - Relative velocity of baryons and dark matter supersonic after recombination.
 - Coherent flows of order $v \sim 30$ km/s on Mpc scales.
- McQuinn & O’Leary (2012):
 - At $z \sim 20$, a 0.3 km/s flow was supersonic, so it needs to be considered.

Shock-heating probably won't destroy the global signal trough



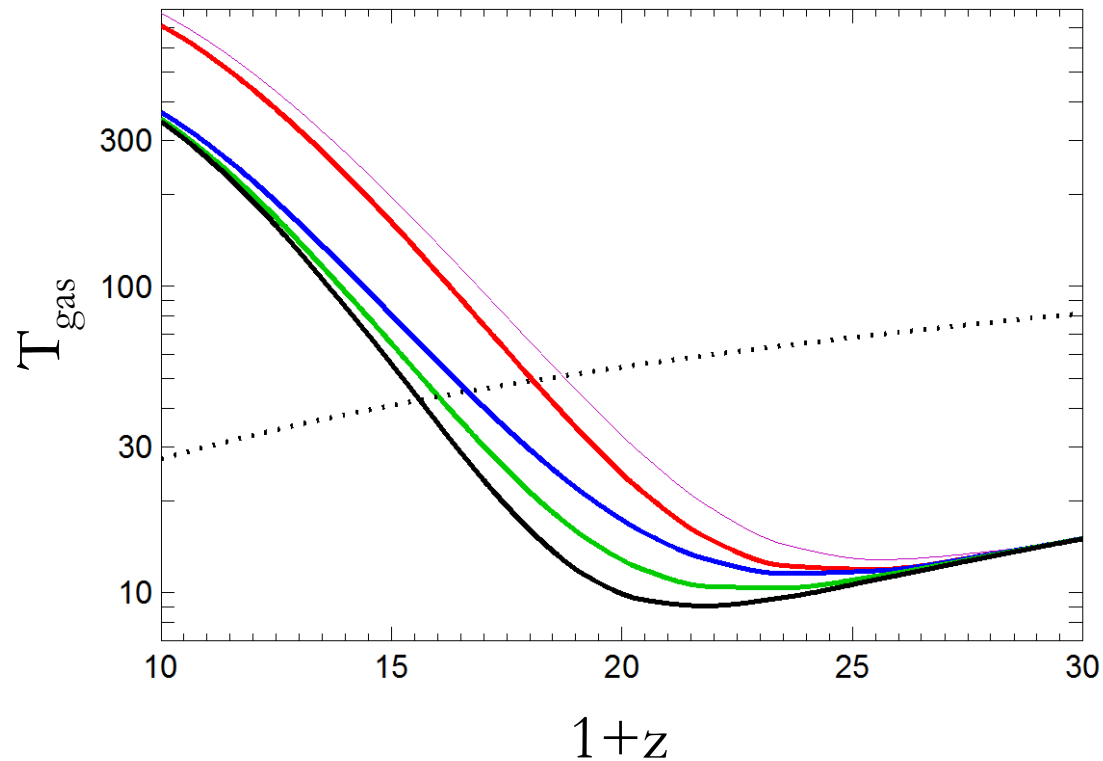
McQuinn & O'Leary (2012)

Shock-heating probably won't destroy the global signal trough



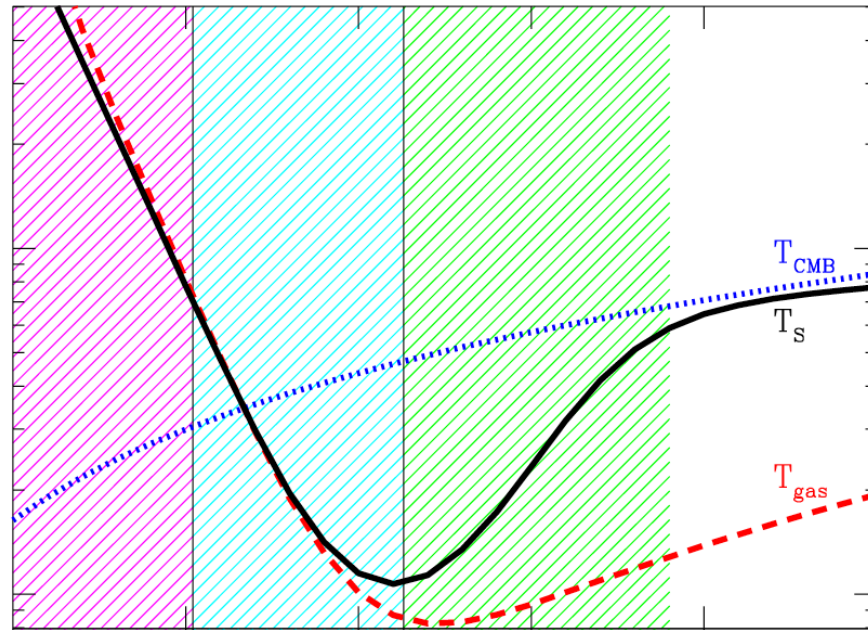
McQuinn & O'Leary (2012)

Now put in star formation, heating,
and feedback as well...



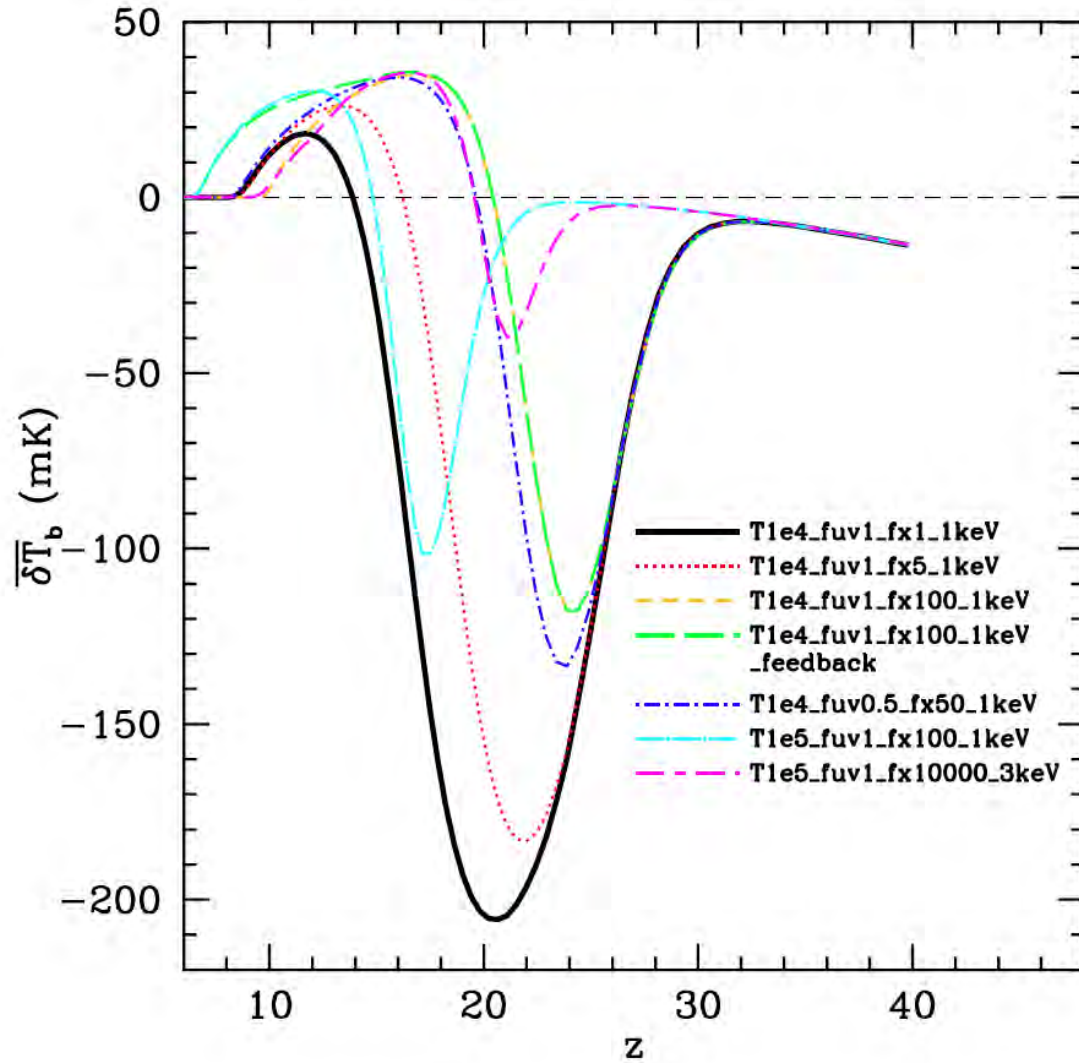
Fialkov et al. (submitted, 2012)

What might go wrong?



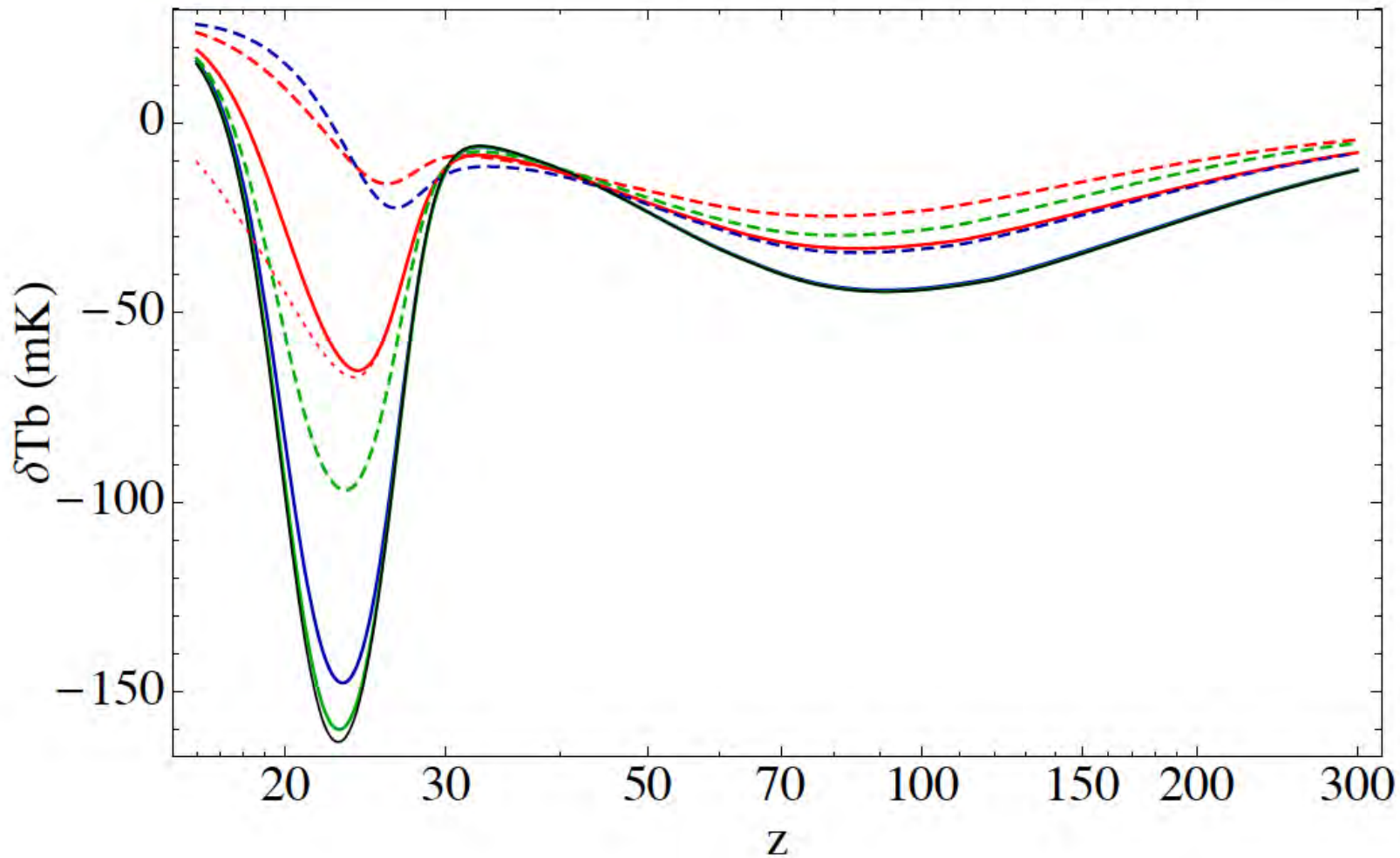
1. IGM physics may cause T_{gas} to deviate from adiabatic cooling.
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Even Xtreme X-rays preserve the trough



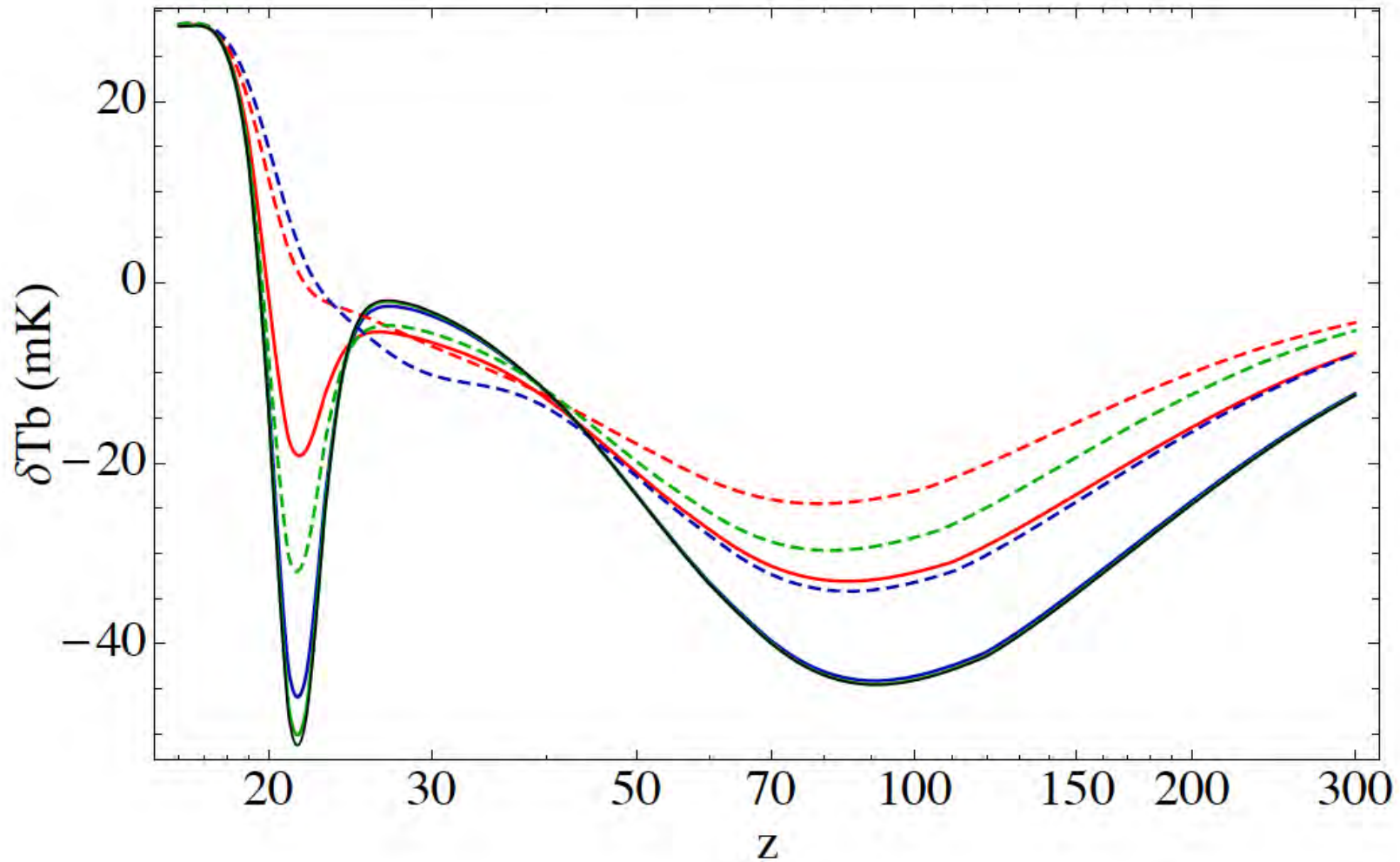
Mesinger, Ferrara, Spiegel (2012)

The trough persists even for (most) models with dark matter annihilations



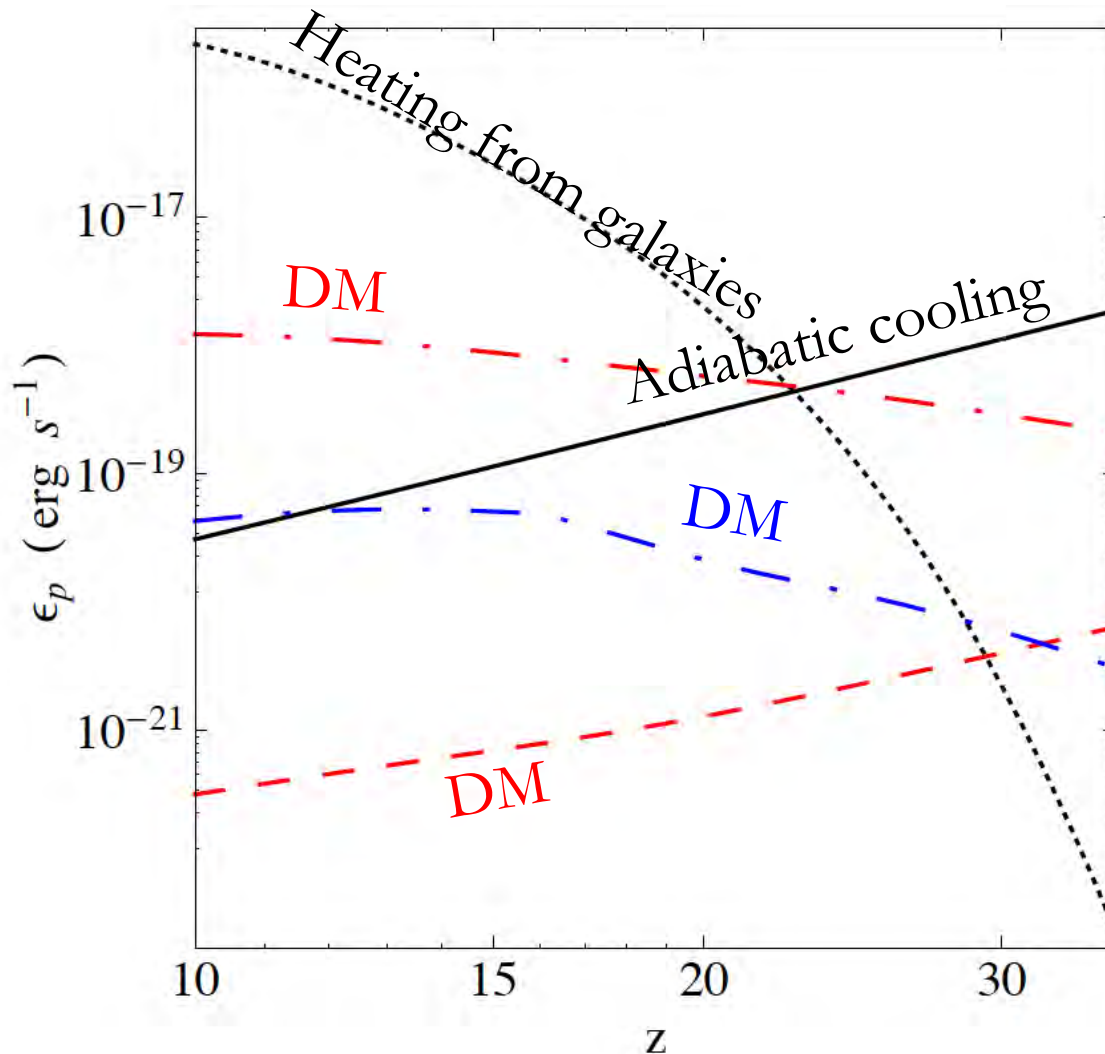
Valdés et al. (2012)

The trough persists even for (most) models with dark matter annihilations



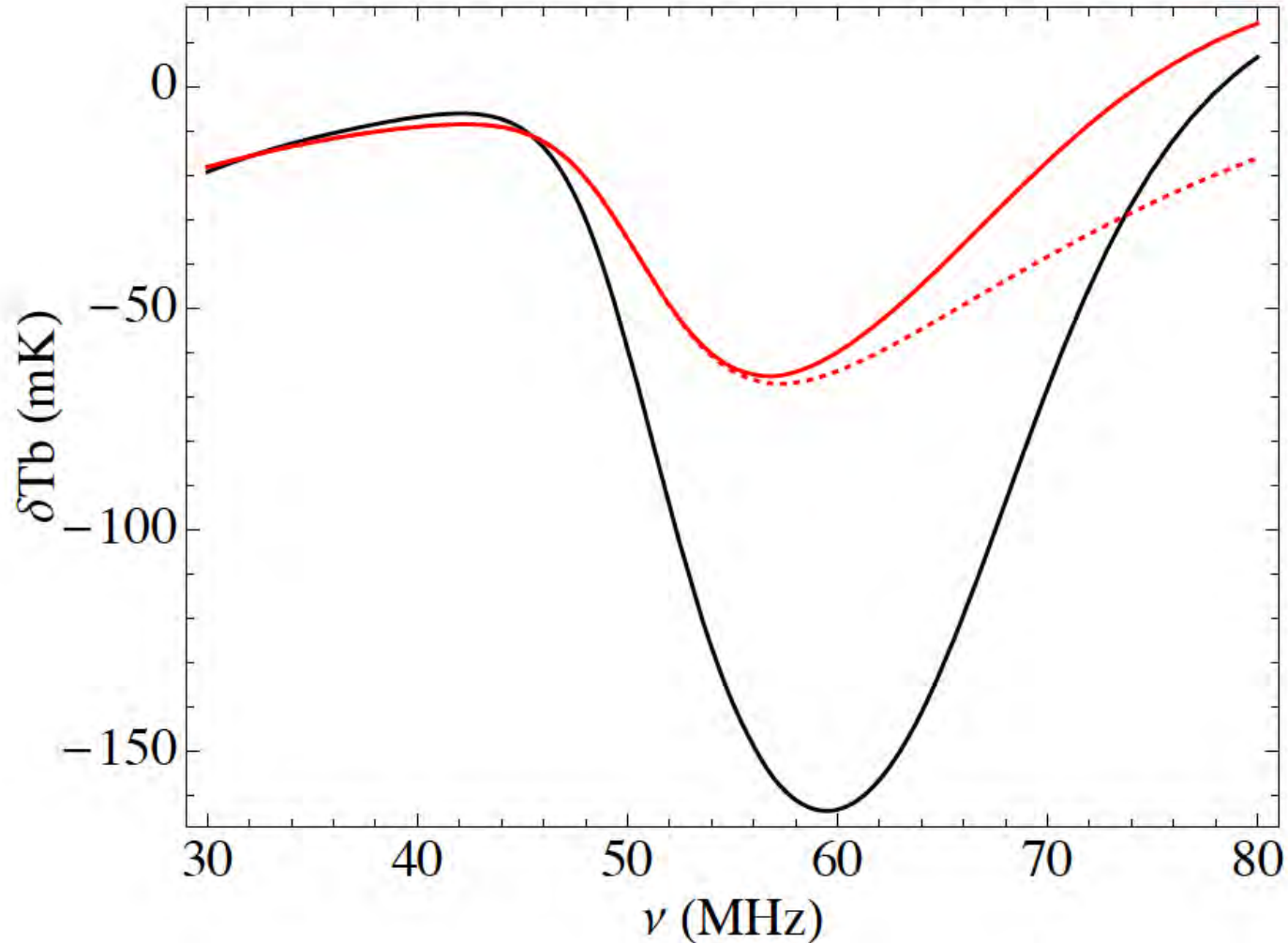
Valdés et al. (2012)

Heating rates are typically different for different sources



Valdés et al. (2012)

The slope of the spectrum may be a useful discriminant



Valdés et al. (2012)

Recent theoretical developments do not change the basic picture, but the considerable uncertainty in theory can result substantial variations.